

Mini-Neutron Star Collision on Laptop

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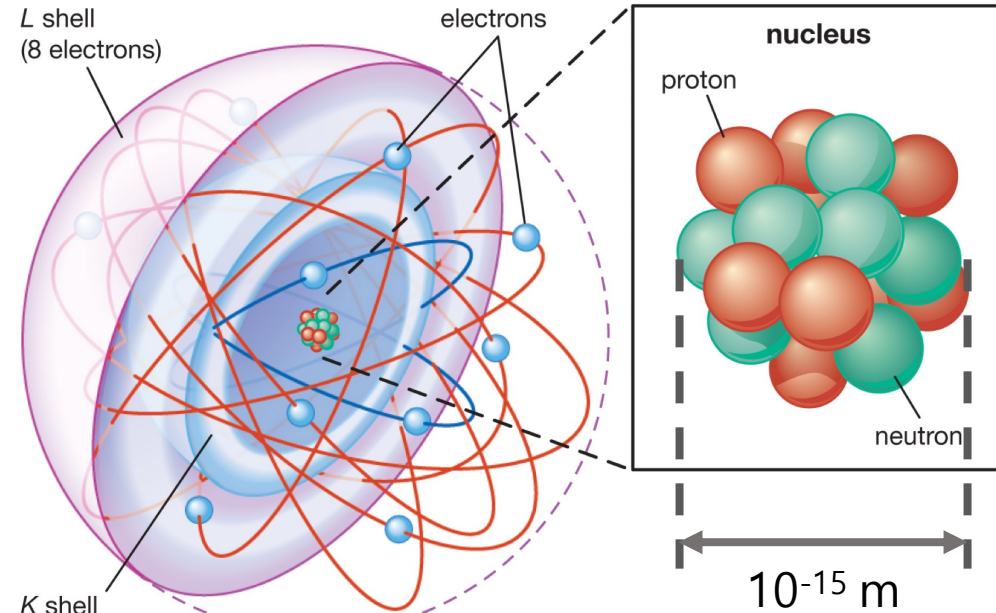
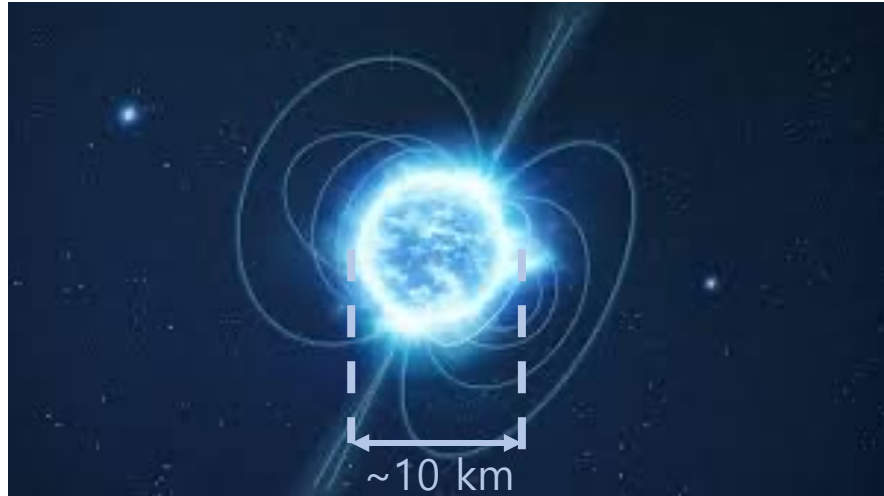


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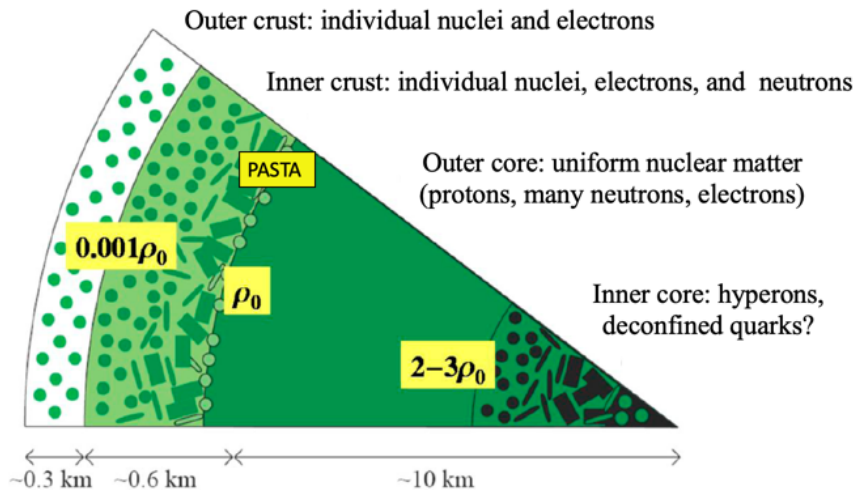
- Introduction
- Transport model
- Comparative study
- Restoring Surface term
- Adopting QMC model
 - Summary

Introduction

Neutron Star, Mini-Neutron Star

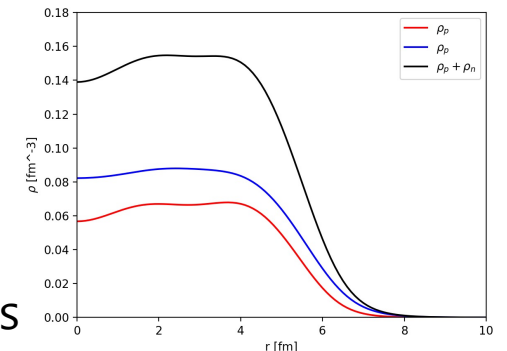


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Mini-Neutron Star = Nucleus
Quantum many-body system
Saturation density ρ_0

Nuclear Physics ~ Astrophysics



Dr. Veronica Dexhemimer, Kent State University

From Nuclei

Binding energy of Nuclei

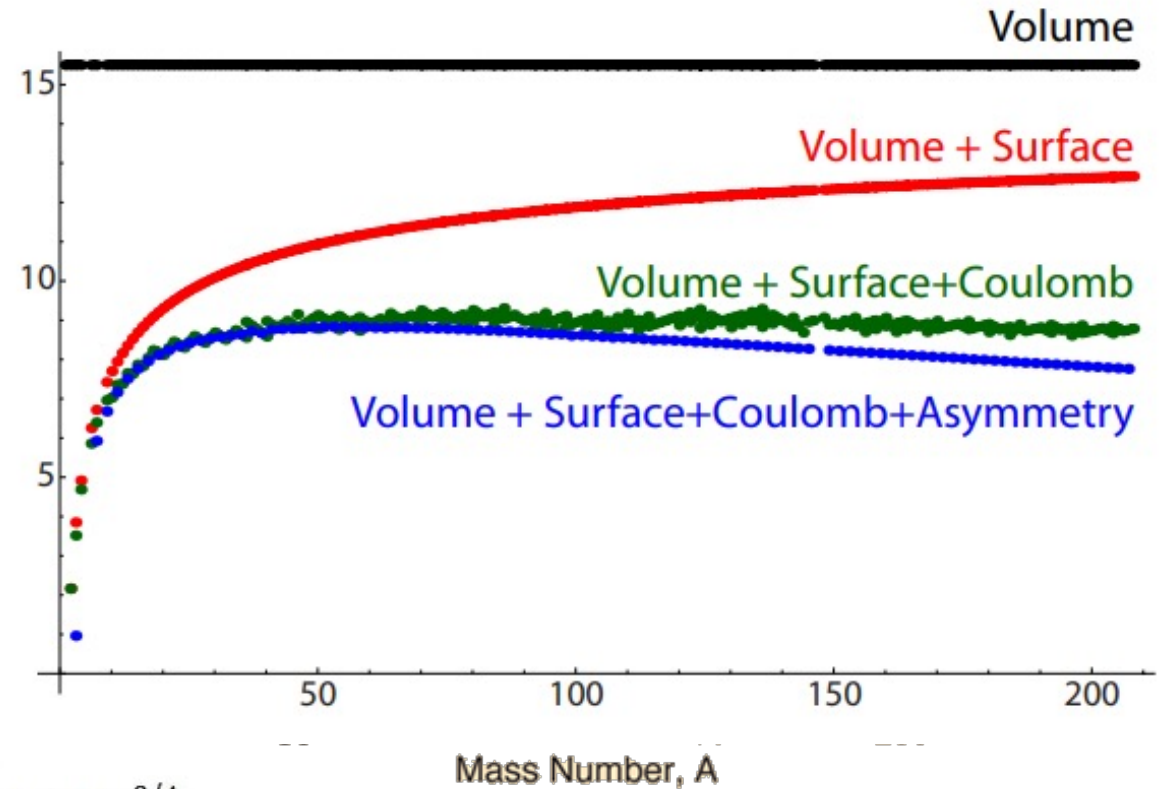
$$m_{\text{nucleus}}c^2 = Zm_p c^2 + Nm_n c^2 - E_b$$

Weizsäcker formula

$$E_b(\text{MeV}) = a_v A - a_s A^{2/3} - a_c \frac{Z^2}{A^{1/3}} - a_A \frac{(A - 2Z)^2}{A} \pm \delta(A, Z)$$

$$\delta(A, Z) = \begin{cases} +\delta_0 & \text{for } Z, N \text{ even} \\ 0 & \\ -\delta_0 & \text{for } Z, N \text{ odd} \end{cases}$$

$$E_b(\text{MeV}) = 15.76A - 17.81A^{2/3} - 0.711 \frac{Z^2}{A^{1/3}} - 23.7 \frac{(N - Z)^2}{A} \pm 34A^{-3/4}$$



From Nuclei

Binding energy of Nuclei

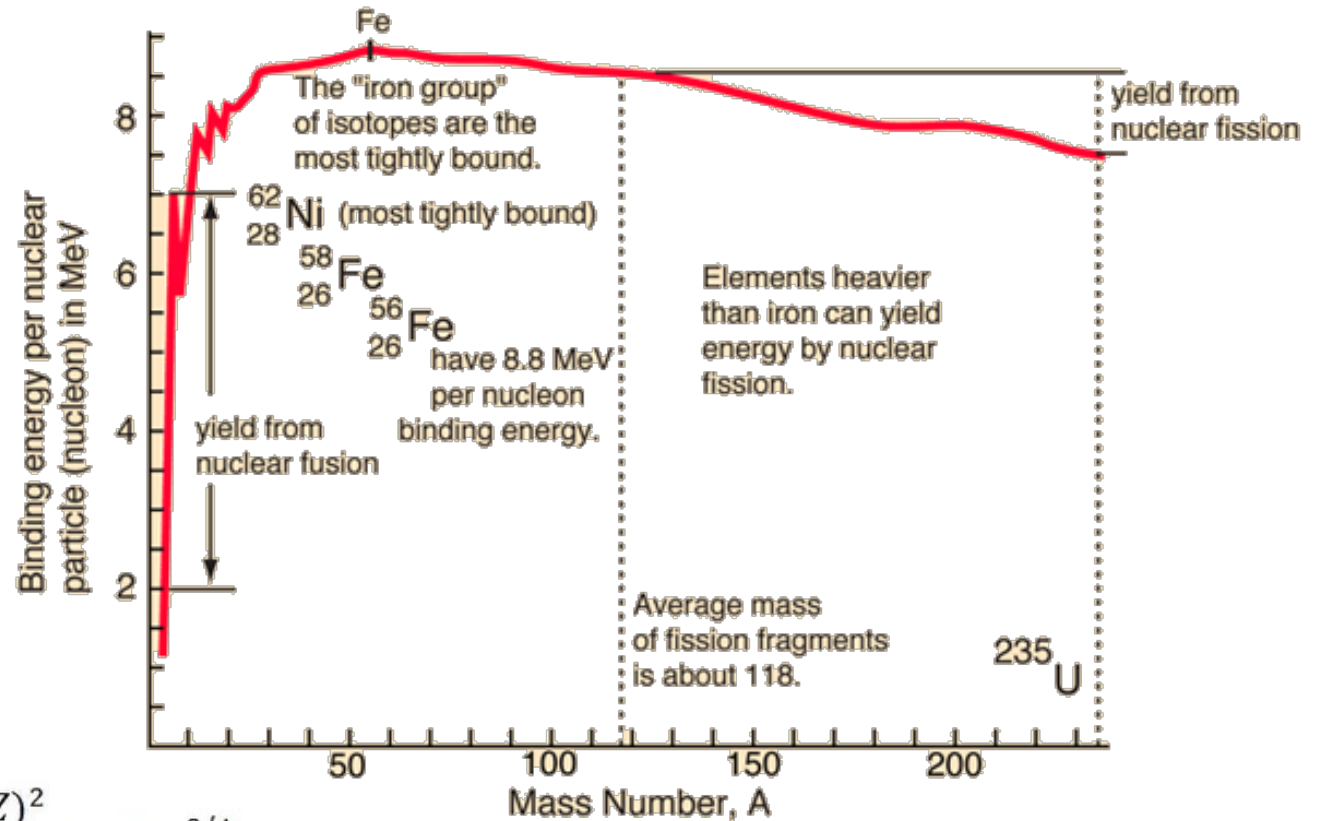
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Nuclear matter properties

Binding energy per Nucleon

$$E(\rho, \alpha) = E_0(\rho) + E_{\text{sym}}(\rho)\alpha^2$$

$$= E_0 + \frac{1}{2!}K_0\chi^2 + \left(S_0 + L\chi + \frac{1}{2!}K_{\text{sym}}\chi^2\right)\alpha^2$$

where,

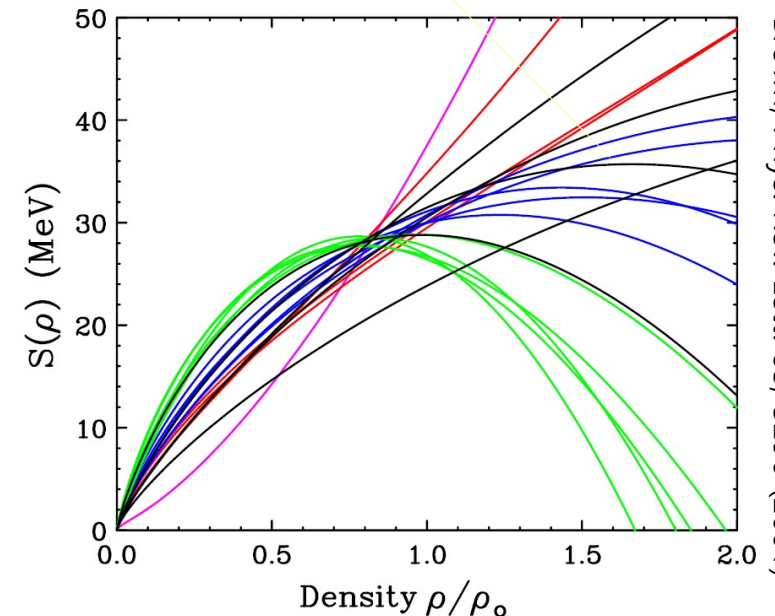
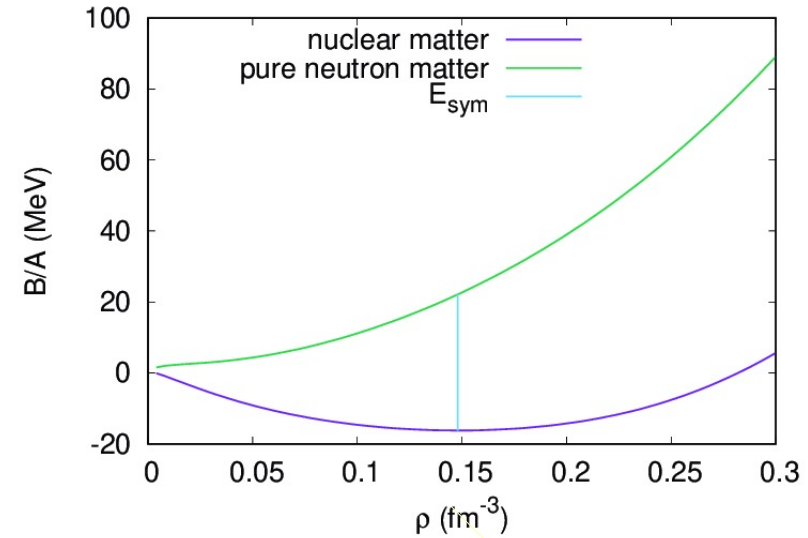
$$\chi = \frac{\rho - \rho_0}{3\rho_0}, \quad \alpha = \frac{\rho_p - \rho_n}{\rho}$$

Incompressibility

$$K_0 = 9\rho_0^2 \left. \frac{d^2 E_0(\rho)}{d\rho^2} \right|_{\rho=\rho_0} = 240 (\pm) \text{ MeV}$$

Slope parameter

$$L = 3\rho_0 \left. \frac{dE_{\text{sym}}(\rho)}{d\rho} \right|_{\rho=\rho_0} = 50 (\pm) \text{ MeV}$$



Brown, Phys. Rev. Lett. 85, 5296 (2001)

Nuclear matter properties

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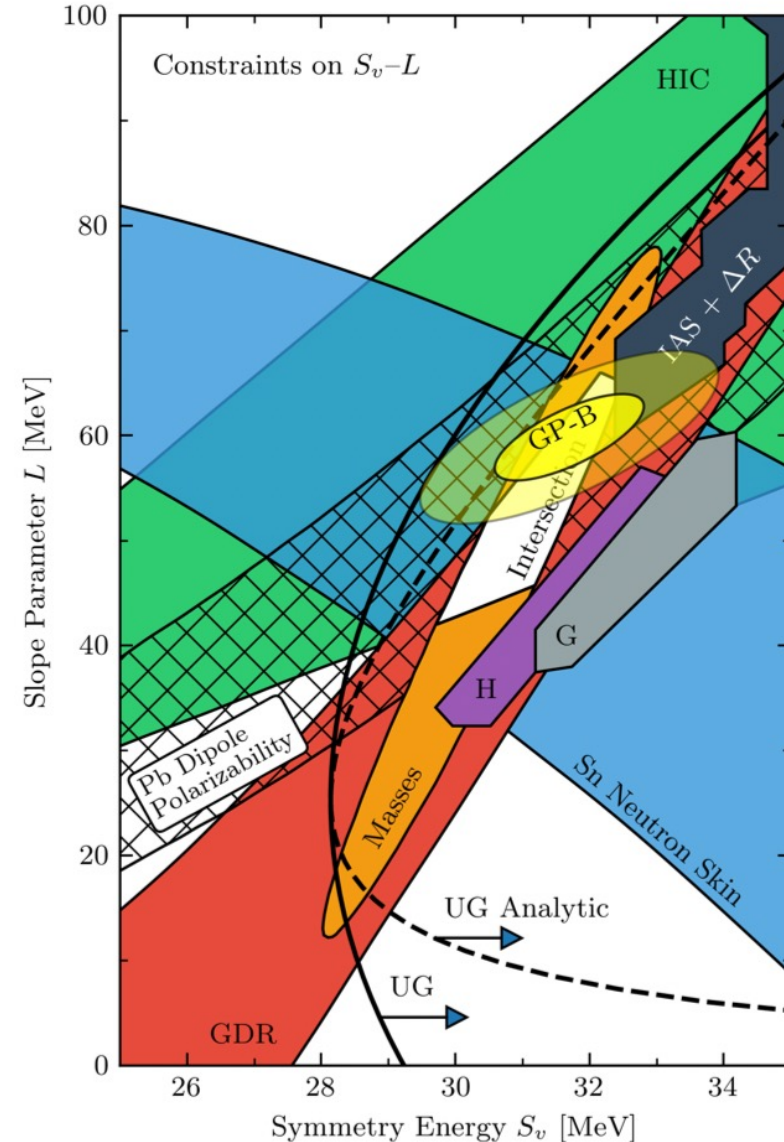
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Bridge between Neutron stars and Nuclei



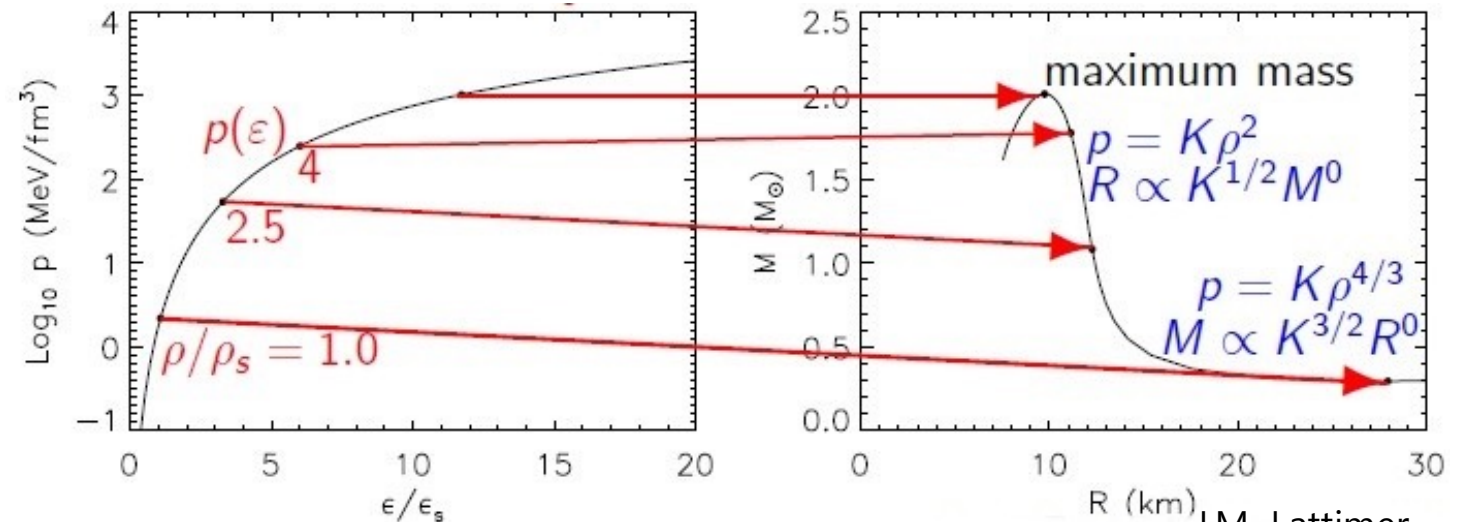
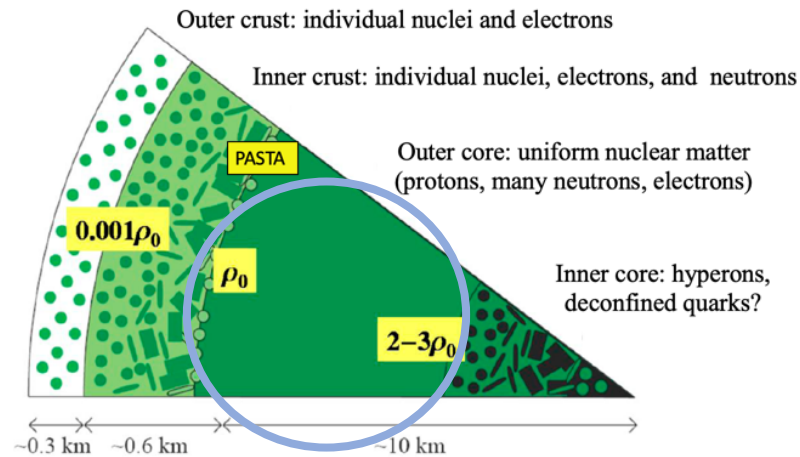
Strong force + Gravity

- Skyrme approach
- RMF approach
- etc...

Tolman–Oppenheimer–Volkoff (TOV) equation

$$\frac{dp}{dr} = -\frac{G}{c^4} \frac{(Mc^2 + 4\pi r^3)(\epsilon + p)}{r(r - 2GM/c^2)}, \quad \frac{dm}{dr} = 4\pi r^2/c^2,$$

Energy density \sim Pressure \rightarrow Radius \sim Mass

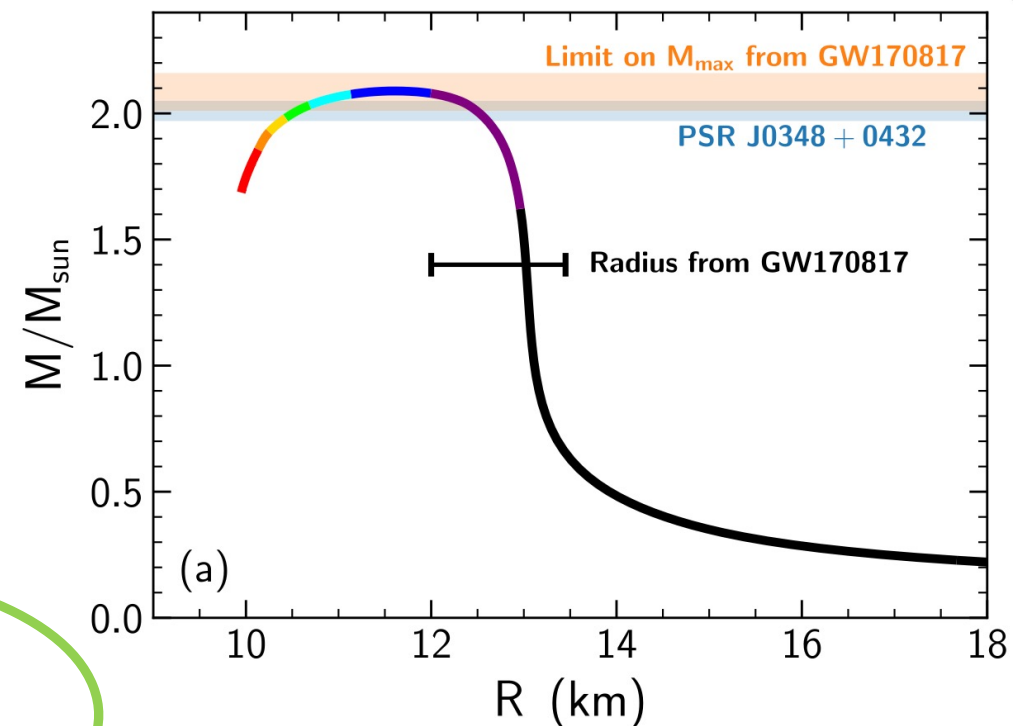
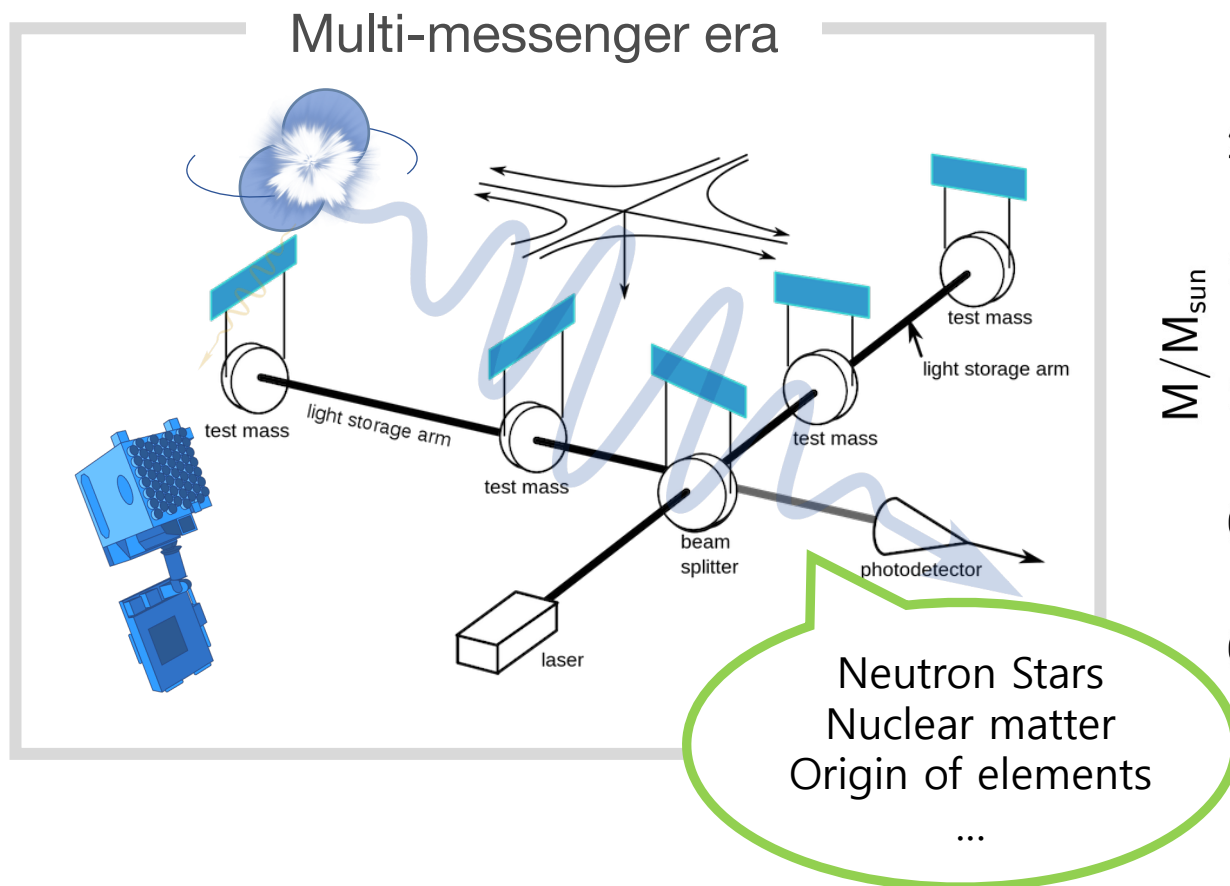


$\rho_0, S_0, K_0, J \dots$

J.M. Lattimer

Neutron Star Collision

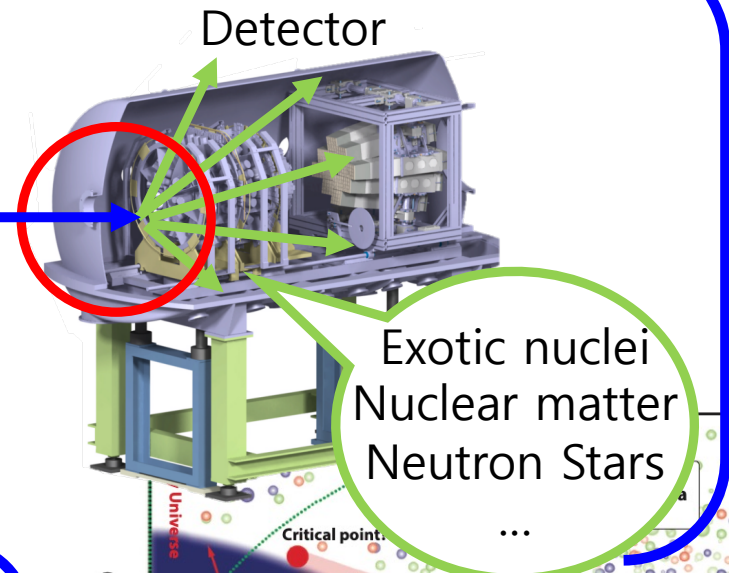
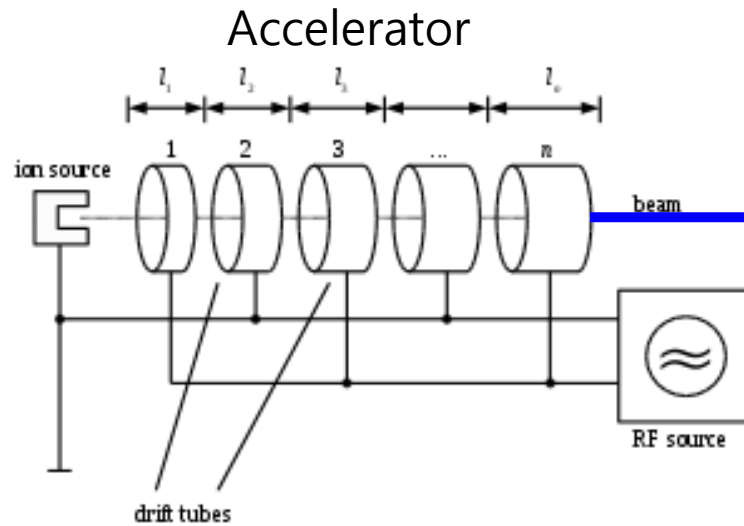
- One of the sources of heavy elements
- And the source of gravitational wave, such as GW170817



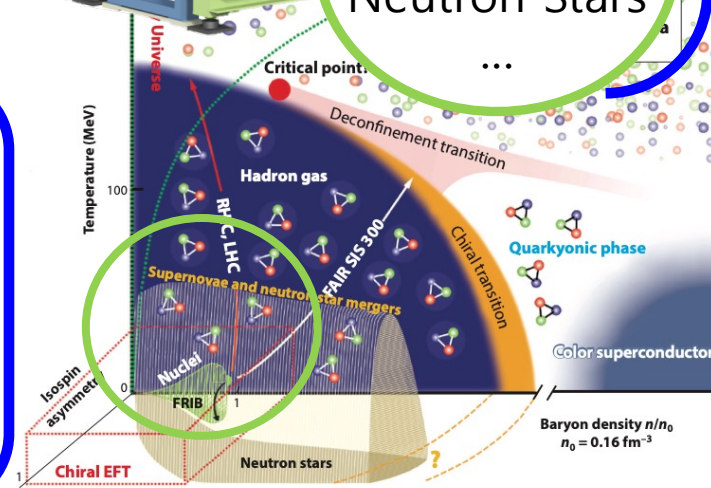
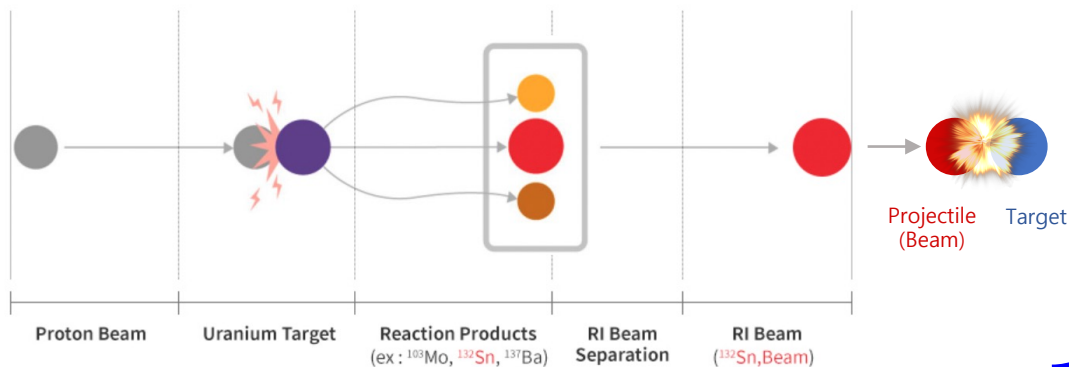
PHYSICAL REVIEW C 101, 034904 (2020)

Mini-Neutron Star Collision on Lab

Heavy Ion Collisions (Experiment) → Information

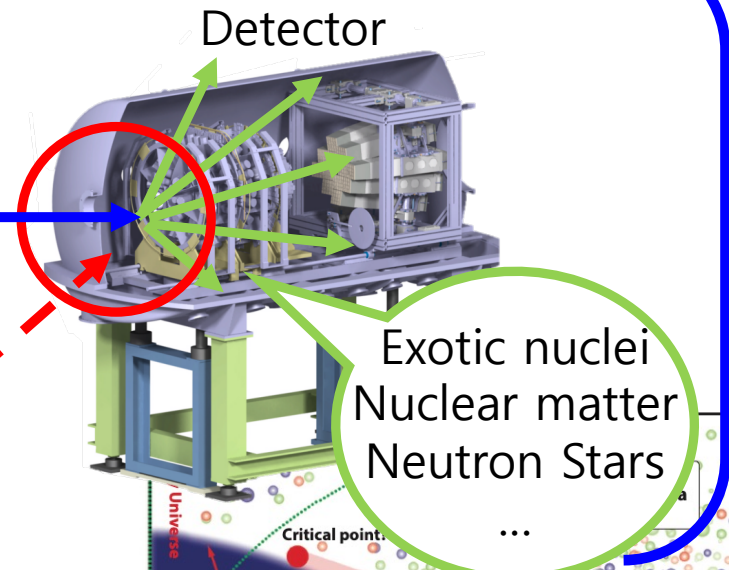
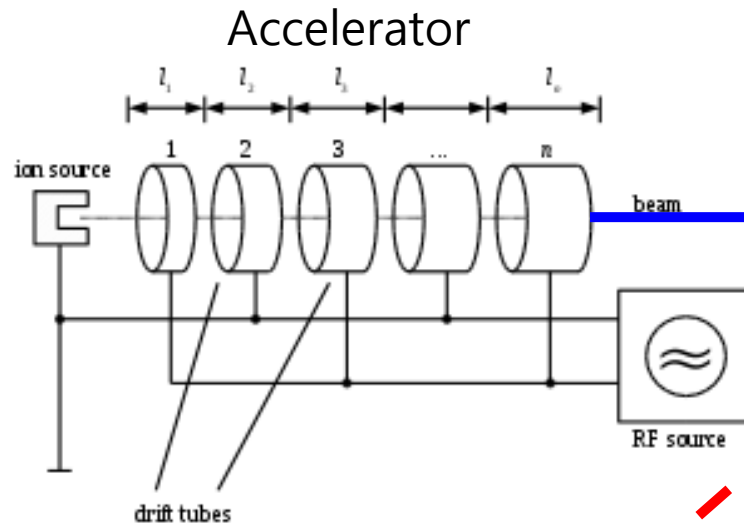


Rare Isotope(RI) beam facilities, FRIB, FAIR & RAON

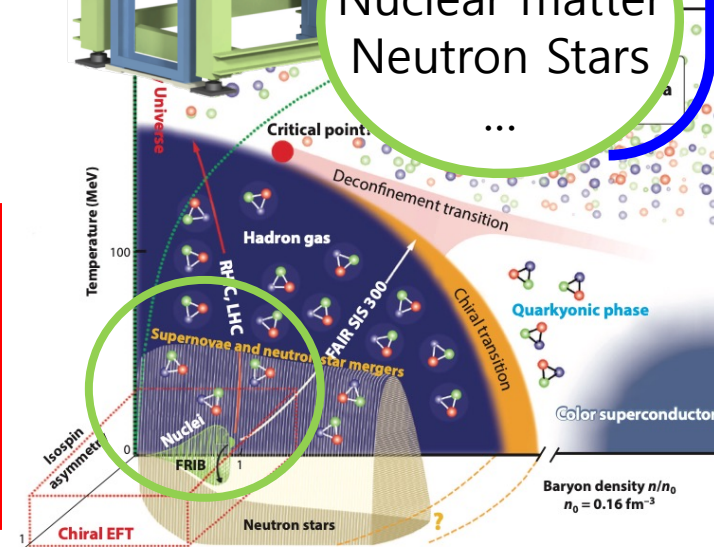


Mini-Neutron Star Collision on Laptop

Heavy Ion Collisions (Experiment) → Information



With numerical model (Theory)



Transport model

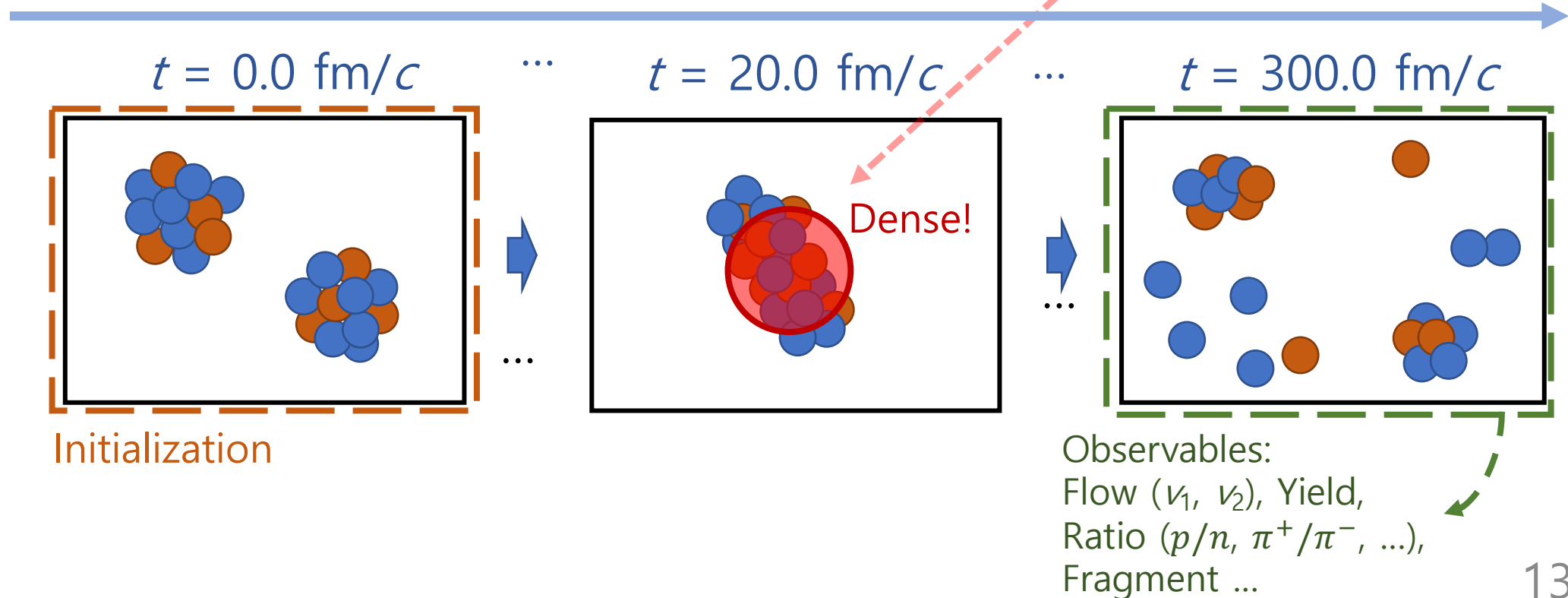
Transport model

Transport model is Theoretical model to allow us study dense matter

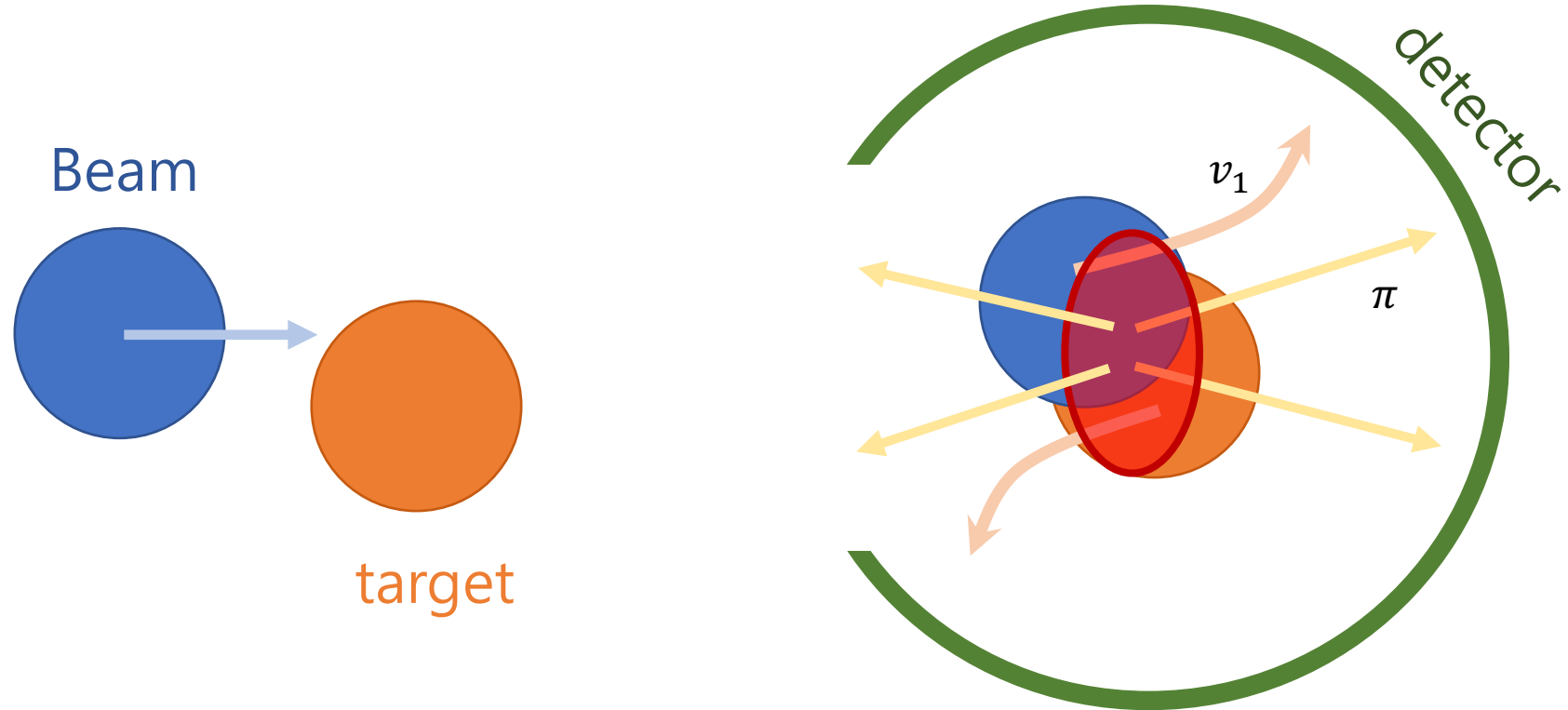
- Semi-classical method
- Hadron degree of freedom

$$\rho_B > \rho_0$$

Full time evolution of Dynamics in Heavy Ion Collision!!



HIC observables



$$\frac{dN}{u_t du_t dy d\phi} = v_0 [1 + 2v_1 \cos(\phi) + 2v_2 \cos(2\phi)],$$
$$v_1 = \left\langle \frac{p_x}{p_t} \right\rangle = \langle \cos(\phi) \rangle, \quad v_2 = \left\langle \left(\frac{p_x}{p_t} \right)^2 - \left(\frac{p_y}{p_t} \right)^2 \right\rangle = \langle \cos(2\phi) \rangle$$

Two type of Transport model

Boltzmann-Uehling-Uhlenbeck (BUU)



$$f(\vec{x}, \vec{p}) = \frac{(2\pi)^3}{N_{TP}} \sum_{i=1}^{AN_{TP}} g_x(\vec{x} - \vec{x}_i) g_p(\vec{p} - \vec{p}_i)$$



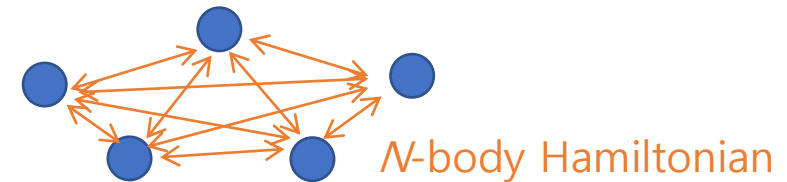
$$\left(\frac{\partial}{\partial t} + \frac{\vec{p}}{m} \cdot \nabla_r - \nabla_r U \cdot \nabla_p \right) f(\vec{r}, \vec{p}; t) = I_{coll}[f; \sigma_{12}]$$

BLOB, GiBUU, pBUU, SMASH and **DJBUU**

Quantum Molecular Dynamics (QMD)



$$f(\vec{x}, \vec{p}) = \exp\left[-\frac{1}{2\sigma_r^2} (\vec{r} - \vec{R})^2 + \left(-\frac{2\sigma_r^2}{\hbar^2}\right) (\vec{p} - \vec{P})^2\right]$$



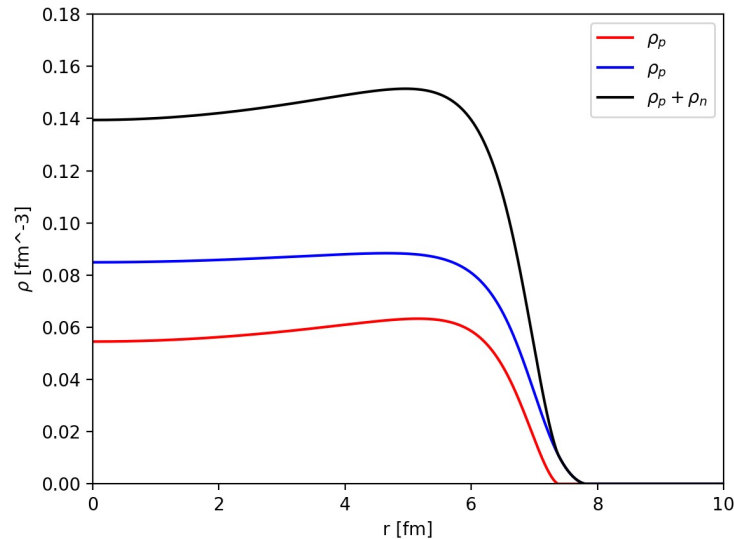
$$H(\vec{r}_n, \vec{p}_n) = \sum_{i=1}^A \frac{\vec{p}_i^2}{2m_i} + \sum_{i<j} V(|\vec{r}_i - \vec{r}_j|)$$

AMD, UrQMD, CoMD, ImQMD and **SQMD** ...

We've developed two Transport model, **DJBUU** and **SQMD**
To study HIC experiments that will be conducted in RAON

Initialization

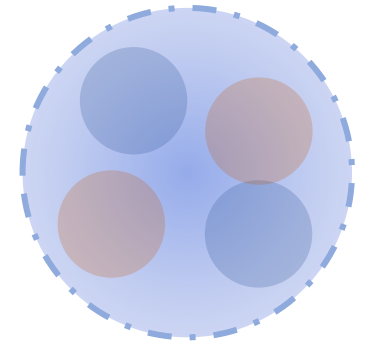
- Wood-Saxon, $\rho(r) = \frac{\rho_0}{1+e^{(r-R)/d}}$
- Relativistic Thomas Fermi (RTF)



Density profile -> Position
Fermi momentum -> momentum

Test particles method,

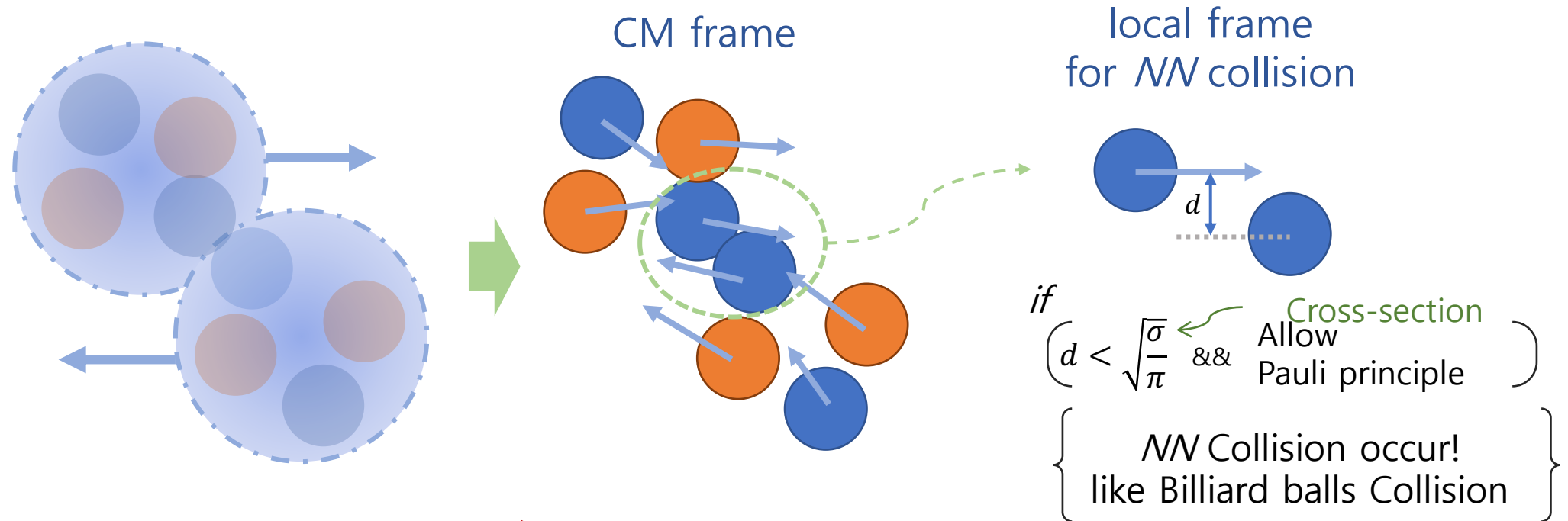
$$g(\vec{u}) = g(u) = N_{m,n} \left[1 - \left(\frac{u}{a_{cut}} \right)^m \right]^n$$



- RCHB (Bubble), DRHBc (deformation)
Density profile of Nucleus → HIC observable, such as v_1

Nucleon-Nucleon collisions

Two Nuclei system \rightarrow Many Nucleons system & Hard collision

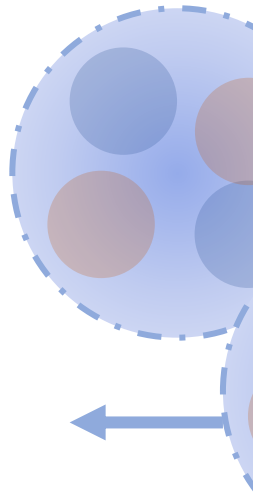


~~NN collisions & free propagation (Cascade mode) not enough~~

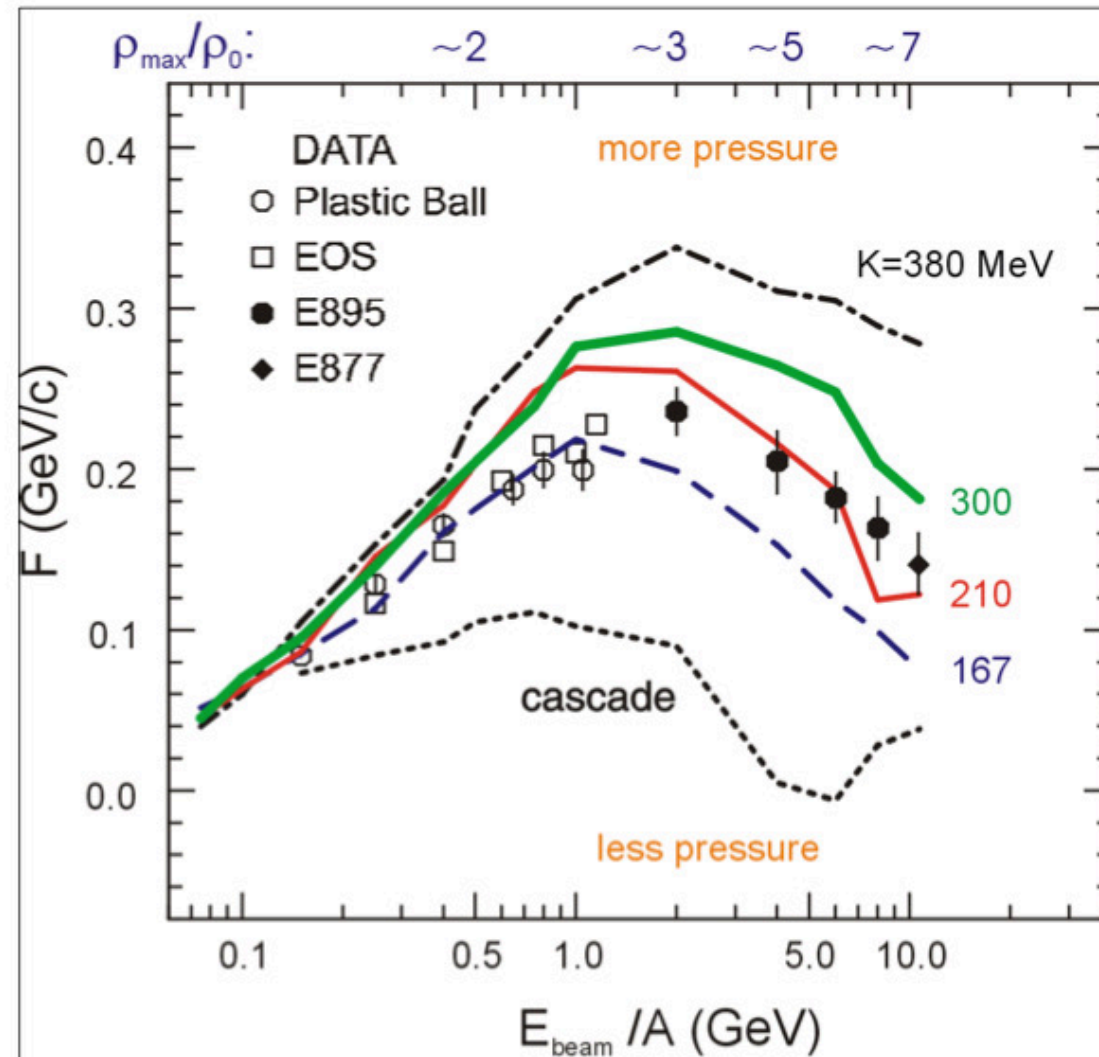
& Propagation with potential

Nucleon-Nucleon collisions

Two Nuclei



NN collisions

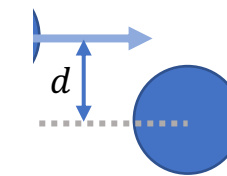


Science 298, 1592 (2002)

d collision

central frame

N collision



Cross-section
&& Allow
Pauli principle

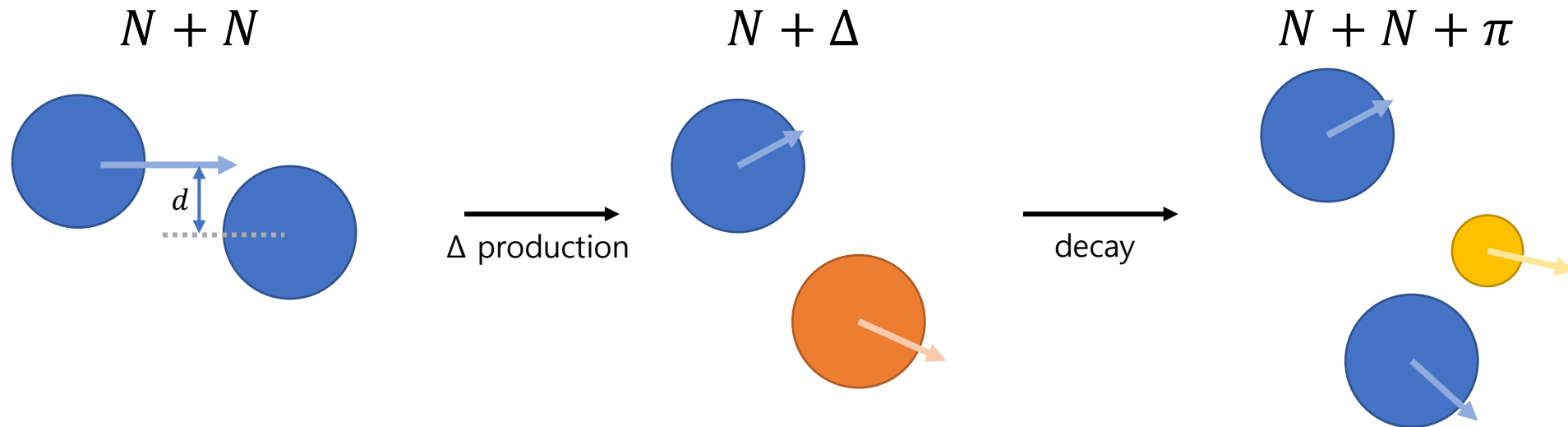
N Collision occur!
billiard balls Collision

not enough

high potential

Nucleon-Nucleon collisions

Pion production in Heavy Ion Collision in intermediate energy reigin

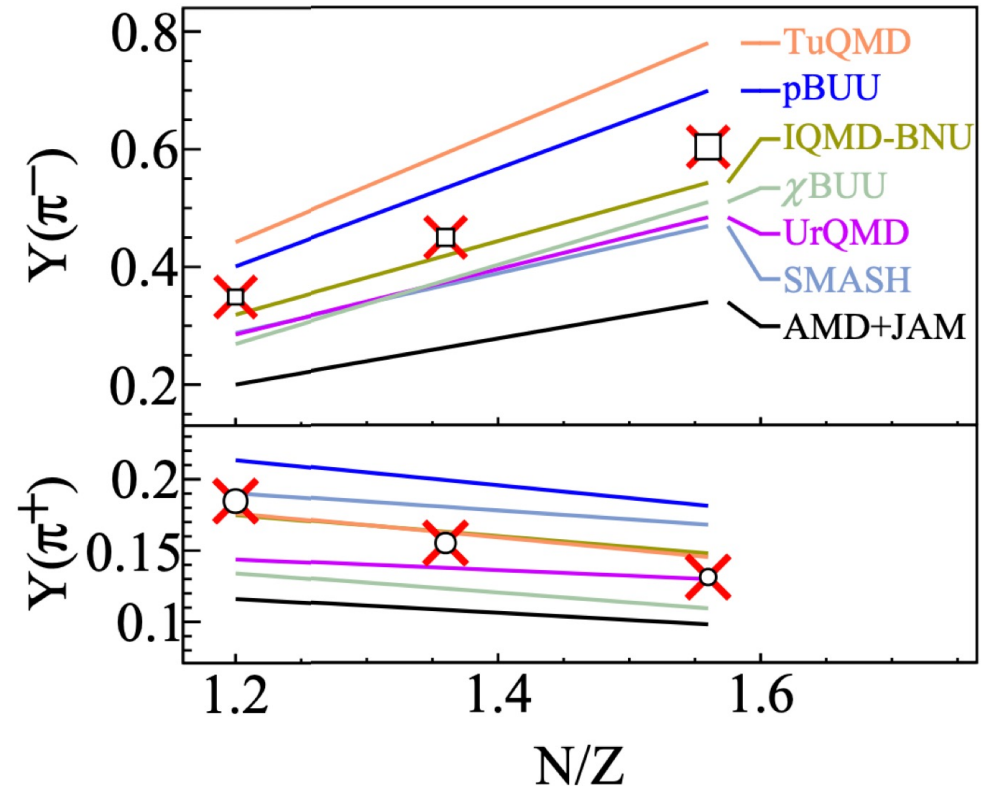
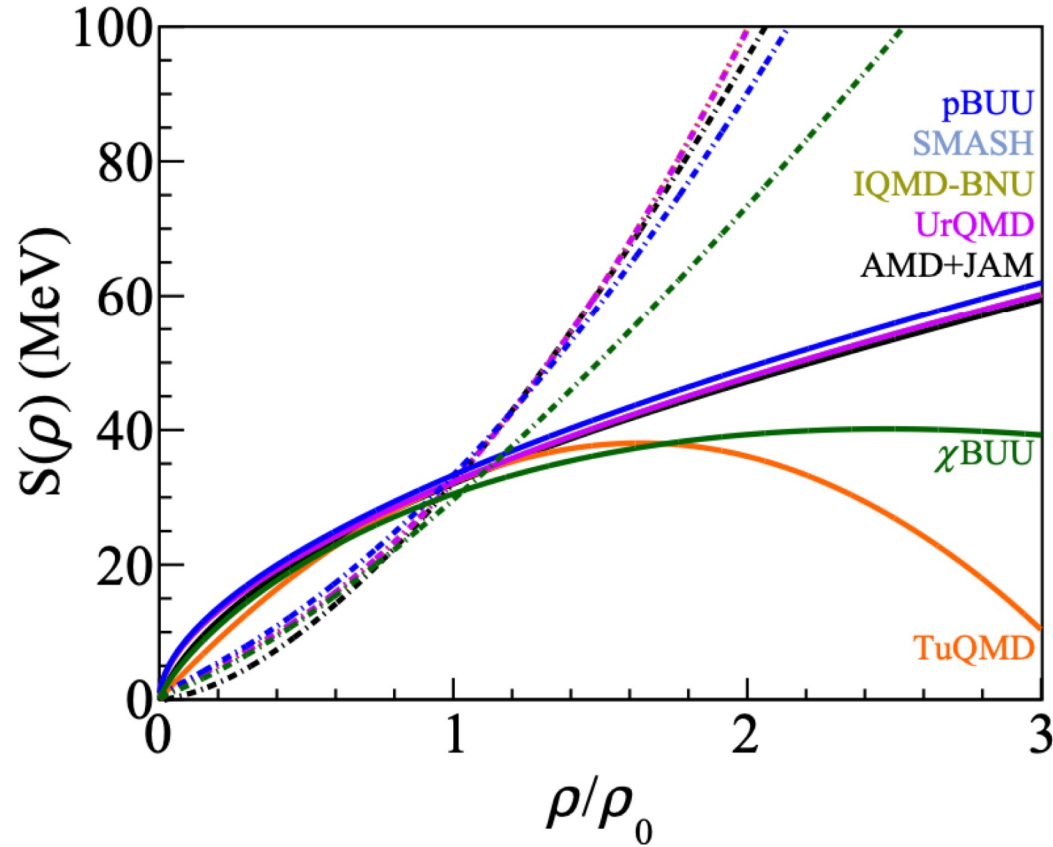


Isospin-dependent Cross-section

$$\sigma_{NN \rightarrow N\Delta}(\rho_B) = \sigma_{NN \rightarrow N\Delta}(0) \times \exp\left(C \frac{\rho_B}{\rho_0}\right) \left(\frac{N}{Z}\right)_{sys}^{x^{\pm,0}}$$

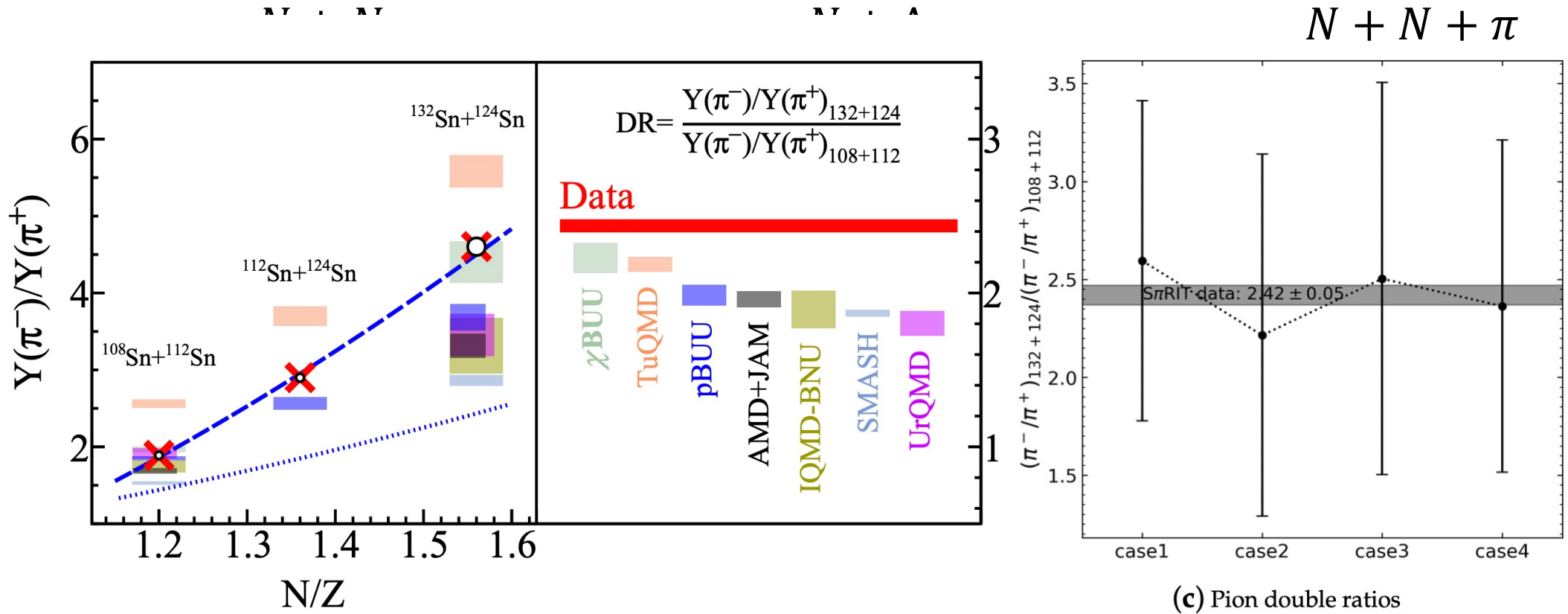
Nucleon-Nucleon collisions

Pion production in Heavy Ion Collision in intermediate energy reigin



Nucleon-Nucleon collisions

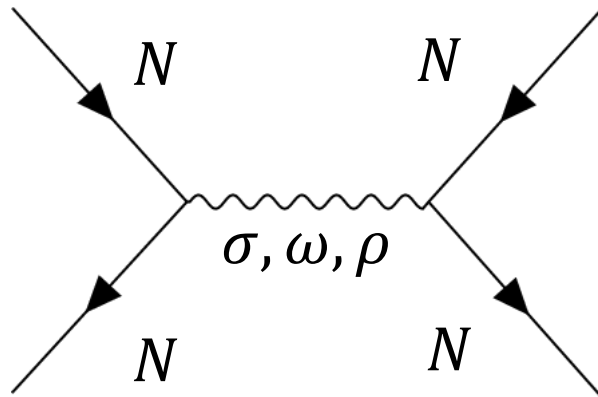
Pion production in Heavy Ion Collision in intermediate energy reigin



Propagation with potential

In DJBUU

Relativistic Mean-Field (RMF) Theory
(ex. QHD, modified Walecka model)



σ : Scalar-Isoscalar - Attractive
 ω : Vector-Isoscalar - Repulsive
 ρ : Vector-Isovector - Repulsive

In SQMD

Non-Relativistic, phenomenological
Skyrme parameterization

$$\begin{aligned}
 U_{tot} = & \frac{\alpha}{2\rho_0} \sum_{i,j \neq i} \rho_{ij} + \frac{\beta}{\gamma + 1} \sum_i \left(\sum_{j \neq i} \frac{\rho_{ij}}{\rho_0} \right)^\gamma \\
 & + \frac{g_{surf}}{2\rho_0} \sum_{i,j \neq i} \nabla_{r_i}^2 (\rho_{ij}) \\
 & + \frac{g_{sym}}{2\rho_0} \sum_{i,j \neq i} [2\delta_{\tau_i \tau_j} - 1] \rho_{ij} \\
 & + \frac{e^2}{2} \sum_{\substack{i,j \neq i, \\ (i,j \text{ for protons})}} \frac{1}{|\vec{r}_i - \vec{r}_j|} \operatorname{erf} \left(\frac{|\vec{r}_i - \vec{r}_j|}{2\sigma_r} \right),
 \end{aligned}$$

Relative Mean Field(RMF) theory

Lagrangian density

$$\begin{aligned} \mathcal{L} = & \bar{\psi} \left[i\gamma_{\mu} \partial^{\mu} - (m_N + g_{\sigma} \sigma) - g_{\omega} \gamma_{\mu} \omega^{\mu} - g_{\rho} \gamma_{\mu} \vec{\tau} \cdot \vec{\rho}^{\mu} - \frac{e}{2} \gamma_{\mu} (1 + \tau^3) A^{\mu} \right] \psi \\ & + \frac{1}{2} (\partial_{\mu} \sigma \partial^{\mu} \sigma - m_{\sigma}^2 \sigma^2) - U(\sigma) + \frac{1}{2} m_{\omega}^2 \omega_{\mu} \omega^{\mu} + \frac{1}{2} m_{\rho}^2 \vec{\rho}_{\mu} \cdot \vec{\rho}^{\mu} \\ & - \frac{1}{4} \Omega_{\mu\nu} \Omega^{\mu\nu} - \frac{1}{4} \vec{R}_{\mu\nu} \vec{R}^{\mu\nu} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \end{aligned}$$

where,

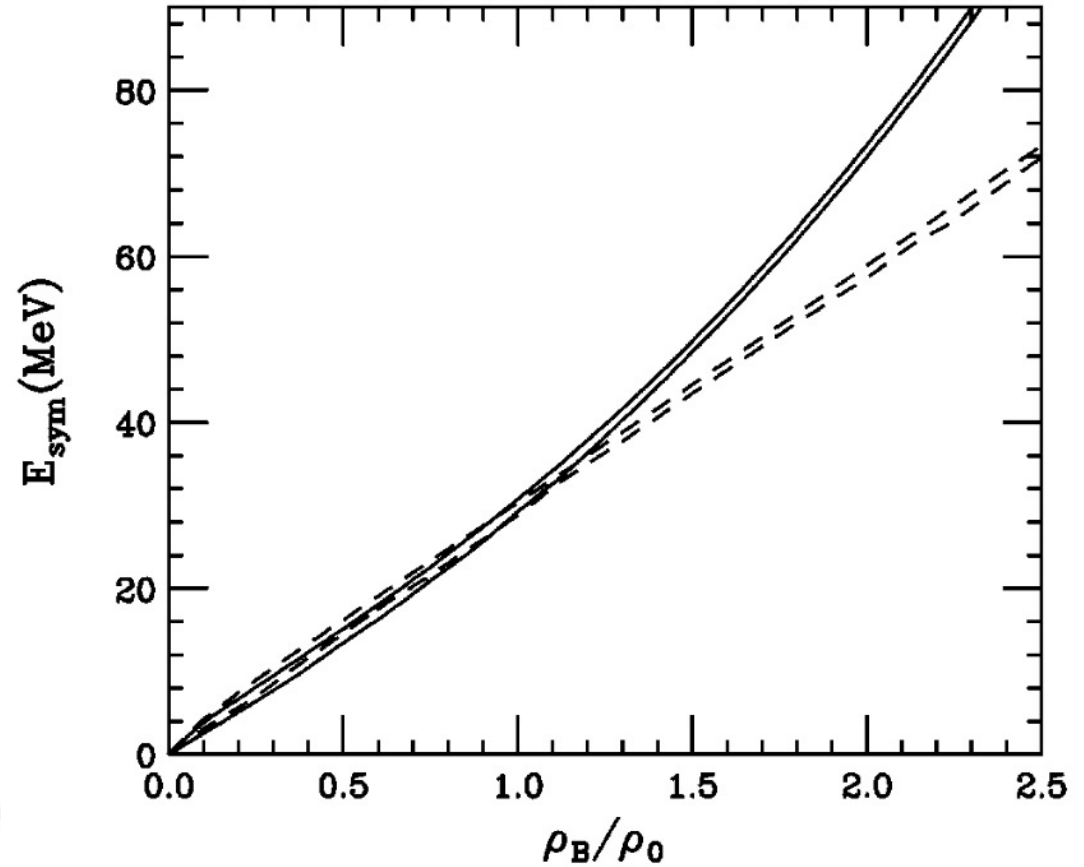
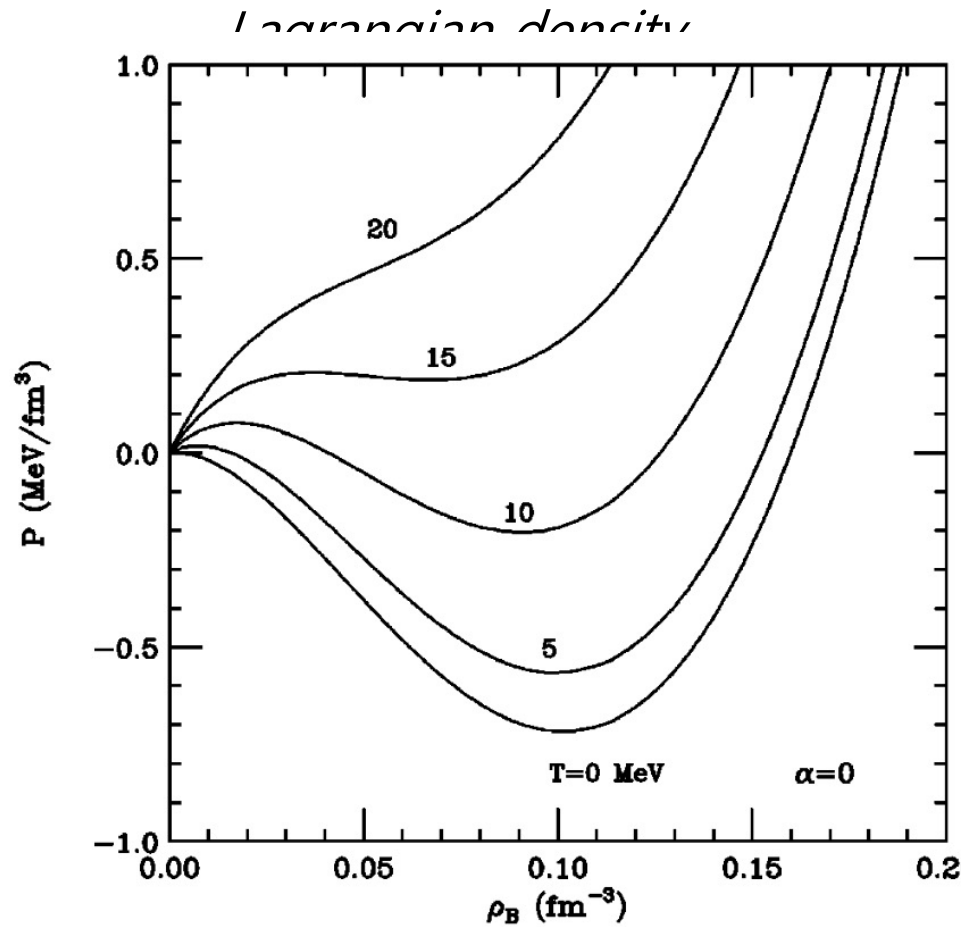
$$U(\sigma) = \frac{1}{3} g_2 \sigma^3 + \frac{1}{4} g_3 \sigma^4, \Omega_{\mu\nu} = \partial_{\mu} \omega_{\nu} - \partial_{\nu} \omega_{\mu}, R_{\mu\nu} = \partial_{\mu} \vec{\rho}_{\nu} - \partial_{\nu} \vec{\rho}_{\mu}, F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu}$$

Table 1. Parameters in the relativistic mean-field model with $f_i = (g_i^2/m_i^2)$ [fm²], $A \equiv a/g_{\sigma}^3$ [fm⁻¹], $B \equiv b/g_{\sigma}^4$, and vacuum masses of nucleon and all mesons in GeV unit.

| f_{σ} | f_{ω} | f_{ρ} | A | B | m_N | m_{σ} | m_{ω} | m_{ρ} |
|--------------|--------------|------------|-------|---------|-------|--------------|--------------|------------|
| 10.33 | 5.42 | 0.95 | 0.033 | -0.0048 | 0.938 | 0.5082 | 0.783 | 0.763 |

PHYSICAL REVIEW C **65** 045201

Relative Mean Field(RMF) theory



PHYSICAL REVIEW C **65** 045201

Relative Mean Field(RMF) theory

To get nuclear potential in DJBUU, We use Relativistic Mean-Field (RMF) Theory and Quantum Hadron Dynamics (QHD), so called Walecka model.

Lagrangian density

$$\begin{aligned}\mathcal{L} = & \bar{\psi} \left[i\gamma_{\mu} \partial^{\mu} - (m_N + g_{\sigma} \sigma) - g_{\omega} \gamma_{\mu} \omega^{\mu} - g_{\rho} \gamma_{\mu} \vec{\tau} \cdot \vec{\rho}^{\mu} - \frac{e}{2} \gamma_{\mu} (1 + \tau^3) A^{\mu} \right] \psi \\ & + \frac{1}{2} (\partial_{\mu} \sigma \partial^{\mu} \sigma - m_{\sigma}^2 \sigma^2) - U(\sigma) + \frac{1}{2} m_{\omega}^2 \omega_{\mu} \omega^{\mu} + \frac{1}{2} m_{\rho}^2 \vec{\rho}_{\mu} \cdot \vec{\rho}^{\mu} \\ & - \frac{1}{4} \Omega_{\mu\nu} \Omega^{\mu\nu} - \frac{1}{4} \vec{R}_{\mu\nu} \vec{R}^{\mu\nu} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}\end{aligned}$$

Euler-Lagrange equation

$$\partial_{\mu} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} q)} - \frac{\partial \mathcal{L}}{\partial q} = 0$$

mean-field approximation

$$\begin{aligned}\sigma & \rightarrow \langle \sigma \rangle \equiv \sigma_0 \\ \omega^{\mu} & \rightarrow \langle \omega^{\mu} \rangle \equiv \delta^{\mu 0} \omega_0 \\ \rho^{\mu} & \rightarrow \langle \vec{\rho}^{\mu} \rangle \equiv \delta^{\mu 0} \vec{\rho}_0\end{aligned}$$

Relative Mean Field(RMF) theory

$$\begin{aligned} \mathcal{L} = & \bar{\psi} \left[i\gamma_\mu \partial^\mu - (m_N + g_\sigma \sigma) - g_\omega \gamma_\mu \omega^\mu - g_\rho \gamma_\mu \vec{\tau} \cdot \vec{\rho}^\mu - \frac{e}{2} \gamma_\mu (1 + \tau^3) A^\mu \right] \psi \\ & + \frac{1}{2} (\partial_\mu \sigma \partial^\mu \sigma - m_\sigma^2 \sigma^2) - U(\sigma) + \frac{1}{2} m_\omega^2 \omega_\mu \omega^\mu + \frac{1}{2} m_\rho^2 \vec{\rho}_\mu \cdot \vec{\rho}^\mu \\ & - \frac{1}{4} \Omega_{\mu\nu} \Omega^{\mu\nu} - \frac{1}{4} \vec{R}_{\mu\nu} \vec{R}^{\mu\nu} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \end{aligned}$$

Euler-Lagrange equation for $\bar{\psi}$

$$\partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \bar{\psi})} - \frac{\partial \mathcal{L}}{\partial \bar{\psi}} = 0 \rightarrow \frac{\partial \mathcal{L}}{\partial \bar{\psi}} = 0$$

$$\left[i\gamma_\mu \partial^\mu - (m_N + g_\sigma \sigma) - g_\omega \gamma_\mu \omega^\mu - g_\rho \gamma_\mu \vec{\tau} \cdot \vec{\rho}^\mu - \frac{e}{2} \gamma_\mu (1 + \tau^3) A^\mu \right] \psi = 0$$

$$(i\gamma_\mu (\partial^\mu - V^\mu) - M^*) \psi = 0$$

Relative Mean Field(RMF) theory

Meson equation

$$\partial_\mu \frac{\partial \mathcal{L}}{\partial(\partial_\mu \sigma)} - \frac{\partial \mathcal{L}}{\partial \sigma} = 0 \rightarrow \partial_\mu \partial^\mu \sigma + m_\sigma^2 \sigma + \frac{\partial U(\sigma)}{\partial \sigma} + g_\sigma \bar{\psi} \psi$$

$$\partial_\mu \sigma \partial^\mu \sigma + m_\sigma^2 \sigma + g_2 \sigma^2 + g_3 \sigma^3 = -g_\sigma \bar{\psi} \psi$$

Mean field approximation: $\sigma \rightarrow \langle \sigma \rangle \equiv \sigma$

$$\partial_\mu \partial^\mu \langle \sigma \rangle + m_\sigma^2 \langle \sigma \rangle + g_2 \langle \sigma \rangle^2 + g_3 \langle \sigma \rangle^3 = -g_\sigma \langle \bar{\psi} \psi \rangle$$

Same way for ω, ρ

$$\partial_\mu \partial^\mu \langle \omega^\nu \rangle + m_\omega^2 \langle \omega^\nu \rangle = g_\omega \langle \bar{\psi} \gamma^\nu \psi \rangle$$

$$\partial_\mu \partial^\mu \langle \rho_3^\nu \rangle + m_\rho^2 \langle \rho_3^\nu \rangle = g_\rho \langle \bar{\psi} \gamma^\nu \tau_3 \psi \rangle$$

Relative Mean Field(RMF) theory

Mean field approximation:

$$\langle \sigma \rangle \equiv \sigma, \quad \langle \omega^\nu \rangle \equiv \omega^0, \quad \langle \rho_3^\nu \rangle \equiv \rho_3^0$$

$$\partial_\mu \partial^\mu \sigma + m_\sigma^2 \sigma + g_2 \sigma^2 + g_3 \sigma^3 = -g_\sigma \rho_s$$

$$\partial_\mu \partial^\mu \omega^0 + m_\sigma^2 \omega^0 = g_\omega j_B^\mu$$

$$\partial_\mu \partial^\mu \rho_3^0 + m_\sigma^2 \rho_3^0 = g_\rho j_{B,I}^\mu$$

where,

$$\langle \bar{\psi} \psi \rangle = \rho_s = \frac{g}{(2\pi)^3} \Sigma \int \frac{m^*}{p^{*0}} dp^3 f(\vec{x}, \vec{p})$$

$$\langle \bar{\psi} \gamma^\nu \psi \rangle = j_B^\mu = \frac{g}{(2\pi)^3} \Sigma \int \frac{p^{*\mu}}{p^{*0}} dp^3 f(\vec{x}, \vec{p})$$

$$\langle \bar{\psi} \gamma^\nu \tau_3 \psi \rangle = j_{B,I}^\mu = \frac{g}{(2\pi)^3} \Sigma \int \frac{p^{*\mu}}{p^{*0}} dp^3 f(\vec{x}, \vec{p})$$

In Uniform nuclear matter, $\partial_\mu \partial^\mu \sigma = \left(\frac{\partial^2}{\partial t^2} - \nabla^2 \right) \sigma = 0$, same for ω, ρ

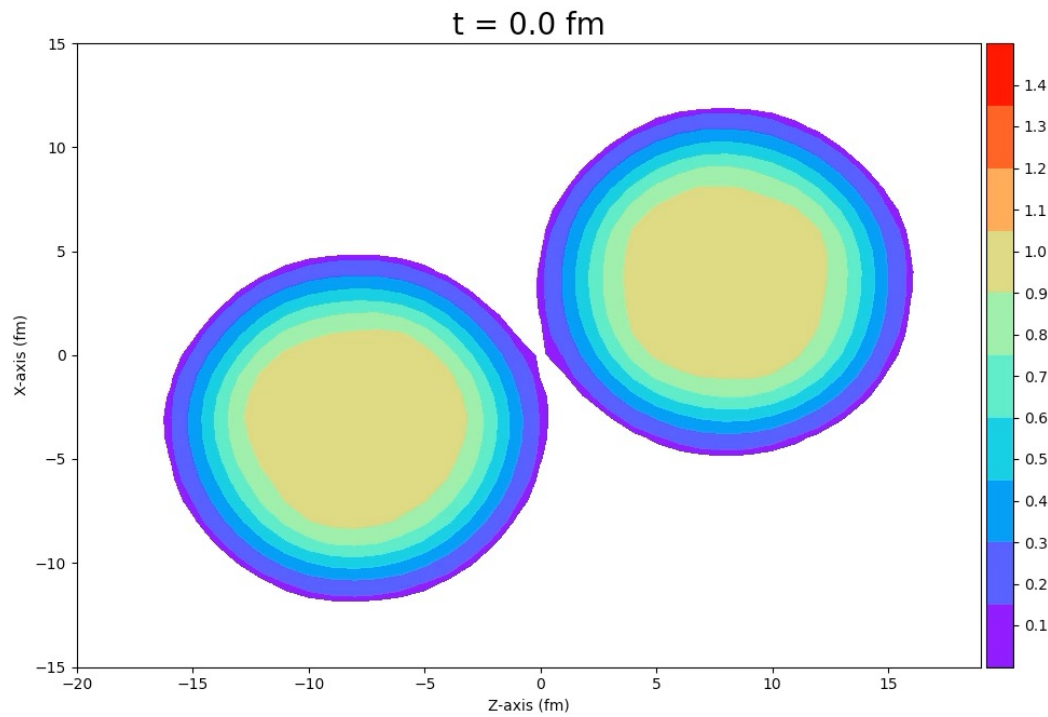
$$m_\sigma^2 \sigma + g_2 \sigma^2 + g_3 \sigma^3 = -g_\sigma \rho_s = -\frac{g_\sigma}{(2\pi)^3} \Sigma \int \frac{m^*}{p^{*0}} dp^3 f(\vec{x}, \vec{p})$$

$$m_\sigma^2 \omega^0 = g_\omega \rho_B = \frac{g_\omega}{(2\pi)^3} \Sigma \int dp^3 f(\vec{x}, \vec{p})$$

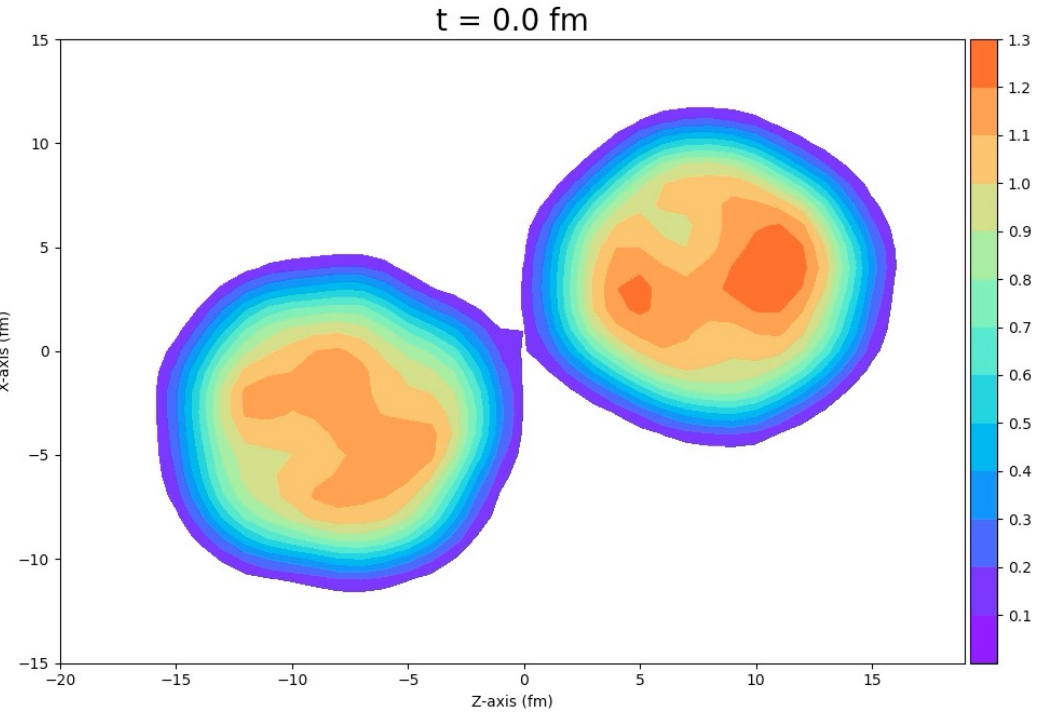
$$m_\sigma^2 \rho_3^0 = g_\rho \rho_{B,I} = \frac{g_\rho}{(2\pi)^3} \Sigma \int dp^3 \tau_3 f(\vec{x}, \vec{p})$$

Full time evolution

$^{197}\text{Au} + ^{197}\text{Au}$, $E_{\text{beam}} = 100 \text{ A MeV}$, $b = 7 \text{ fm}$



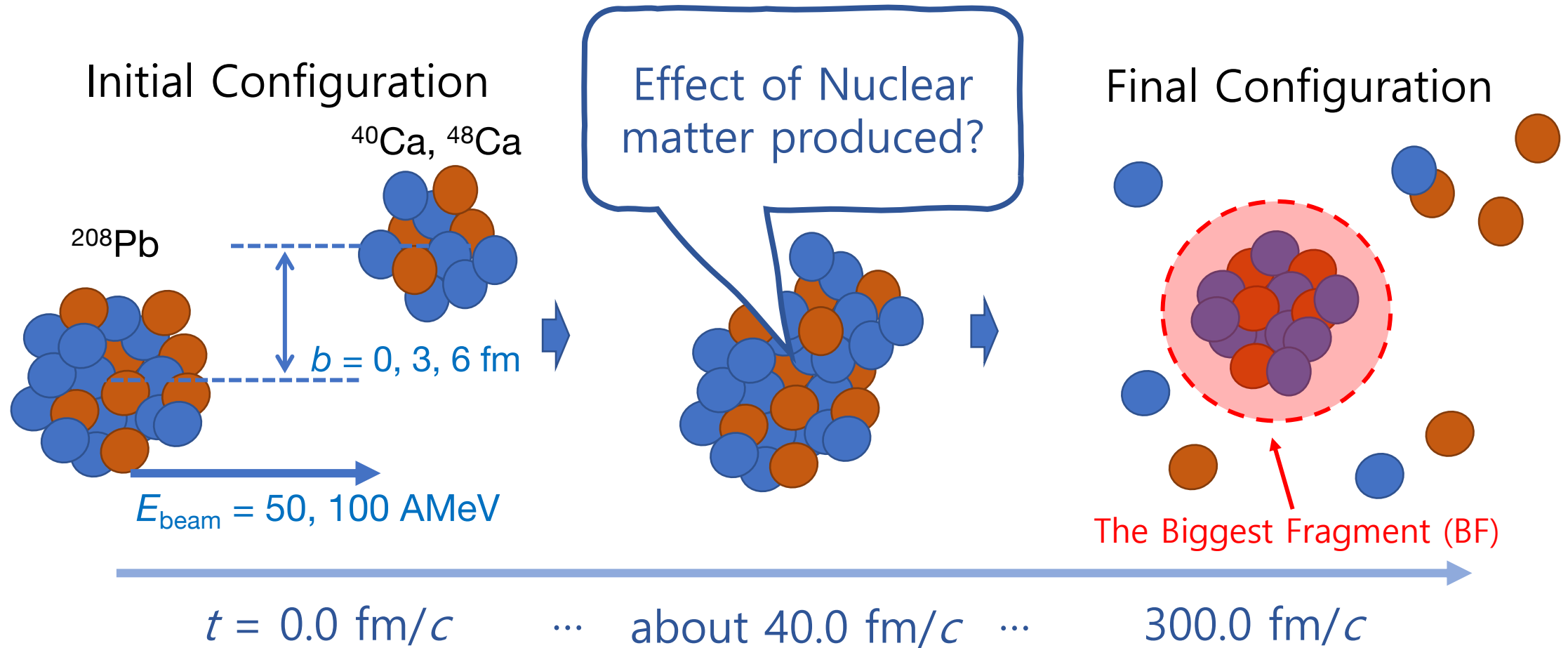
1 event, $N_{TP} = 50$
corresponding to 50 events (DJBUU work)



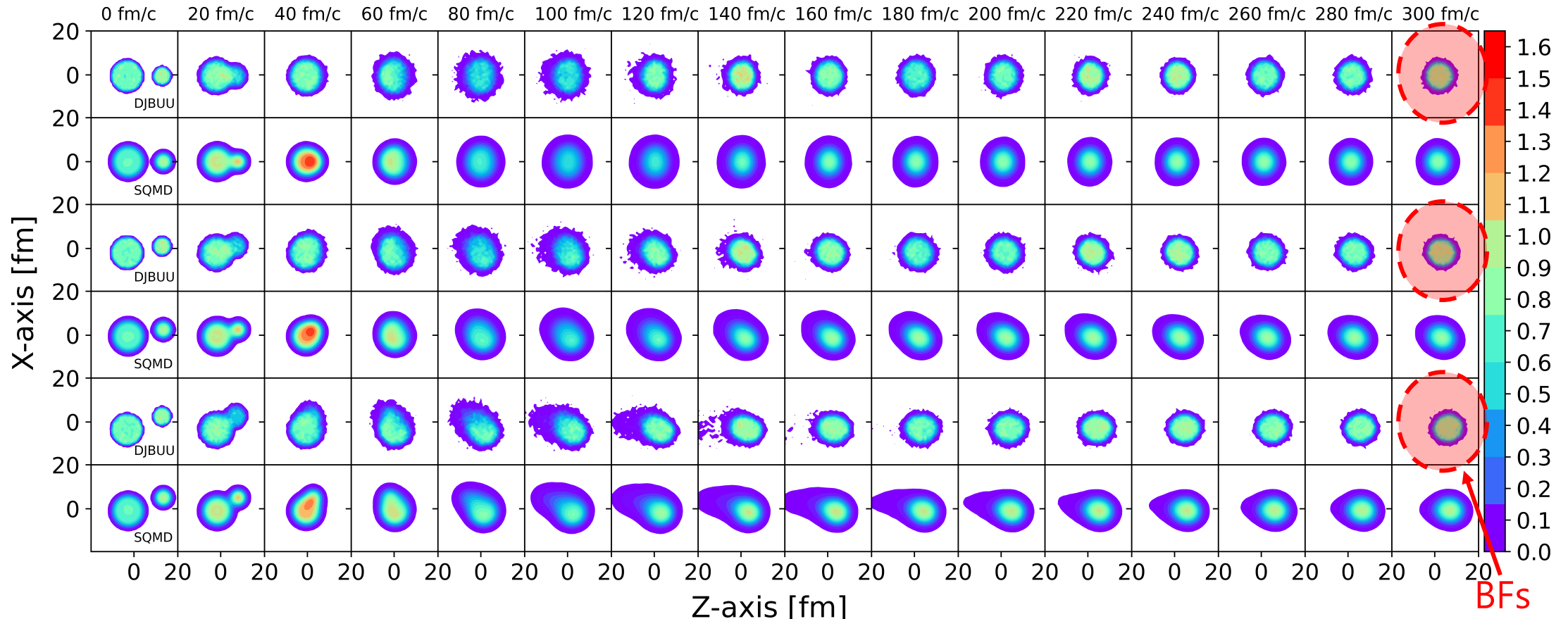
1 event (SQMD work)

Comparative study

Comparative study : Fragment

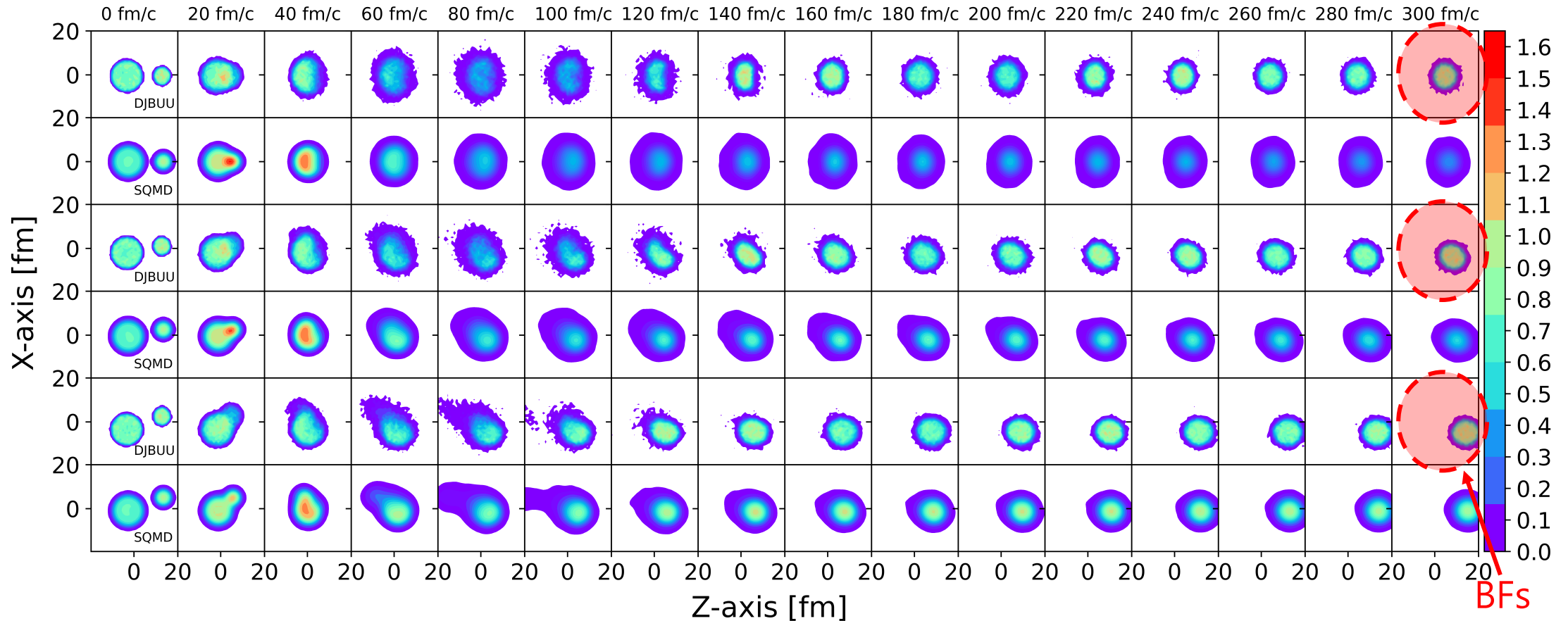


Comparative study : Fragment



Density distribution in the collision plane. For comparison, the results of DJBUU and SQMD are shown alternatively. From top to bottom, the systems are 208 Pb + 40 Ca at E beam = 50 A MeV

Comparative study : Fragment



Density distribution in the collision plane. For comparison, the results of DJBUU and SQMD are shown alternatively. From top to bottom, the systems are 208 Pb + 40 Ca at E beam = 100 A MeV.

Comparative study : Fragment

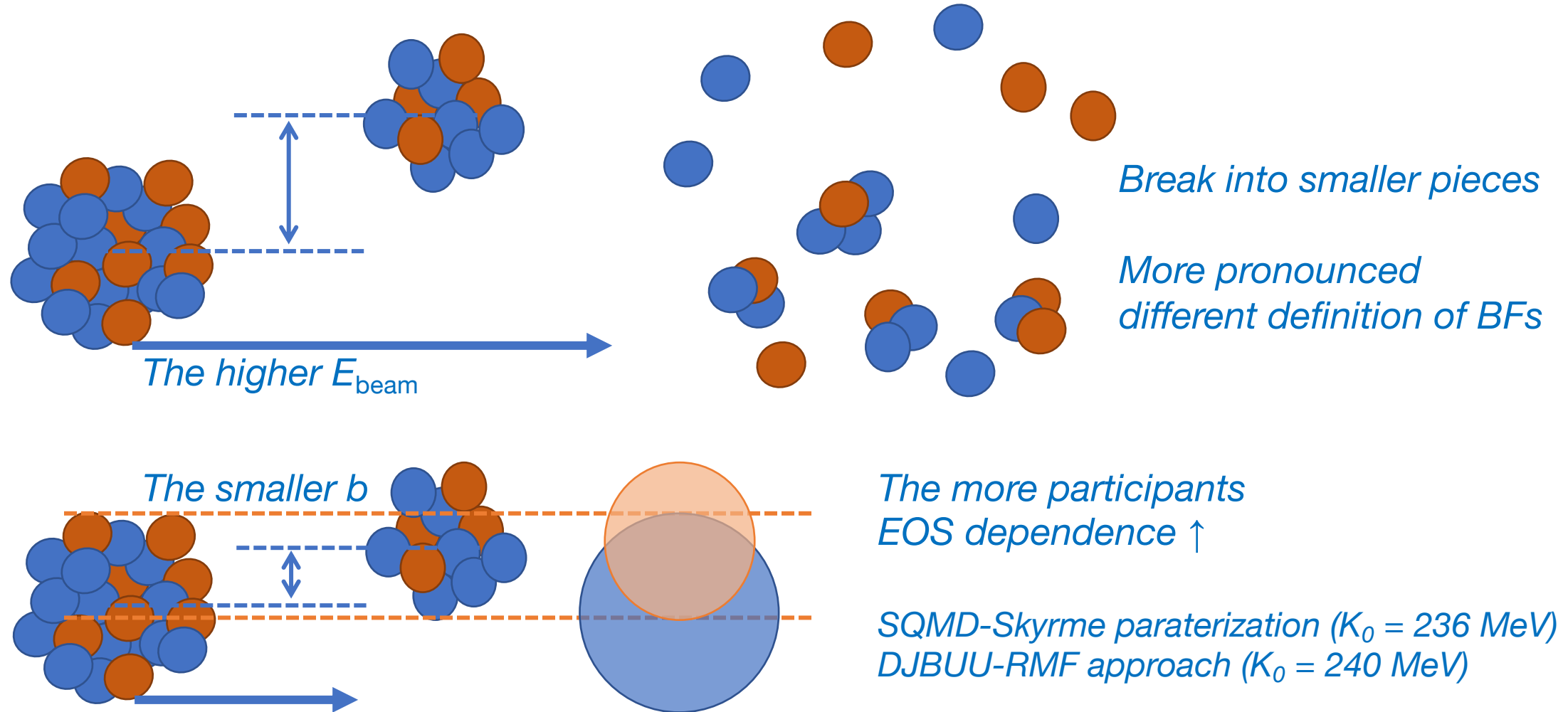
| Target | E_{beam} (AMeV) | b (fm) | DJBUU | SQMD |
|------------------|--------------------------|----------|--|--|
| ^{40}Ca | 50 | 0 | $^{163}_{73}\text{Ta}$, $^{162}_{73}\text{Ta}$, $^{164}_{73}\text{Ta}$, $^{163}_{74}\text{W}$ | $^{163}_{69}\text{Tm}$, $^{173}_{74}\text{W}$, $^{169}_{72}\text{Hf}$ |
| | | 3 | $^{163}_{73}\text{Ta}$, $^{165}_{74}\text{W}$, $^{164}_{73}\text{Ta}$, | $^{169}_{72}\text{Hf}$, $^{173}_{74}\text{W}$, $^{172}_{74}\text{W}$ |
| | | 6 | $^{167}_{74}\text{W}$, $^{169}_{75}\text{Re}$, $^{165}_{73}\text{Ta}$, $^{168}_{75}\text{Re}$ | $^{168}_{72}\text{Hf}$, $^{164}_{70}\text{Yb}$, $^{169}_{72}\text{Hf}$ |
| | 100 | 0 | $^{123}_{56}\text{Ba}$, $^{121}_{55}\text{Cs}$, $^{124}_{57}\text{La}$, $^{122}_{56}\text{Ba}$, $^{124}_{56}\text{Ba}$ | $^{78}_{33}\text{As}$, $^{114}_{50}\text{Sn}$, $^{124}_{54}\text{Xe}$ |
| | | 3 | $^{130}_{59}\text{Pr}$, $^{130}_{58}\text{Ce}$, $^{128}_{57}\text{La}$, $^{128}_{58}\text{Ce}$, $^{129}_{58}\text{Ce}$, $^{127}_{58}\text{Ce}$, $^{127}_{57}\text{La}$ | $^{125}_{53}\text{I}$, $^{128}_{56}\text{Ba}$, $^{132}_{57}\text{La}$ |
| | | 6 | $^{145}_{64}\text{Gd}$, $^{144}_{64}\text{Gd}$, $^{146}_{65}\text{Tb}$, $^{147}_{65}\text{Tb}$ | $^{151}_{64}\text{Gd}$, $^{149}_{63}\text{Eu}$, $^{154}_{66}\text{Dy}$ |
| ^{48}Ca | 50 | 0 | $^{161}_{72}\text{Hf}$, $^{162}_{72}\text{Hf}$, $^{160}_{71}\text{Lu}$, $^{159}_{71}\text{Lu}$ | $^{167}_{70}\text{Yb}$, $^{167}_{71}\text{Lu}$, $^{170}_{71}\text{Lu}$ |
| | | 3 | $^{162}_{72}\text{Hf}$, $^{164}_{73}\text{Ta}$ | $^{165}_{70}\text{Yb}$, $^{167}_{70}\text{Yb}$, $^{167}_{71}\text{Lu}$ |
| | | 6 | $^{164}_{72}\text{Hf}$, $^{163}_{72}\text{Hf}$, $^{166}_{73}\text{Ta}$, $^{165}_{72}\text{Hf}$ | $^{165}_{69}\text{Tm}$, $^{159}_{68}\text{Er}$, $^{164}_{69}\text{Tm}$ |
| | 100 | 0 | $^{113}_{51}\text{Sb}$, $^{115}_{52}\text{Te}$, $^{114}_{51}\text{Sb}$, $^{116}_{52}\text{Te}$, $^{112}_{51}\text{Sb}$ | $^{58}_{25}\text{Mn}$, $^{74}_{32}\text{Ge}$, $^{107}_{48}\text{Pd}$ |
| | | 3 | $^{121}_{54}\text{Xe}$, $^{122}_{55}\text{Cs}$, $^{120}_{54}\text{Xe}$, $^{123}_{55}\text{Cs}$, $^{121}_{55}\text{Cs}$ | $^{120}_{52}\text{Te}$, $^{106}_{45}\text{Rh}$, $^{113}_{48}\text{Cd}$ |
| | | 6 | $^{140}_{62}\text{Sm}$, $^{139}_{62}\text{Sm}$, $^{138}_{61}\text{Pm}$, $^{137}_{61}\text{Pm}$, $^{137}_{60}\text{Nd}$ | $^{147}_{62}\text{Sm}$, $^{153}_{64}\text{Gd}$, $^{148}_{62}\text{Sm}$ |

The higher beam energy

The smaller impact parameter

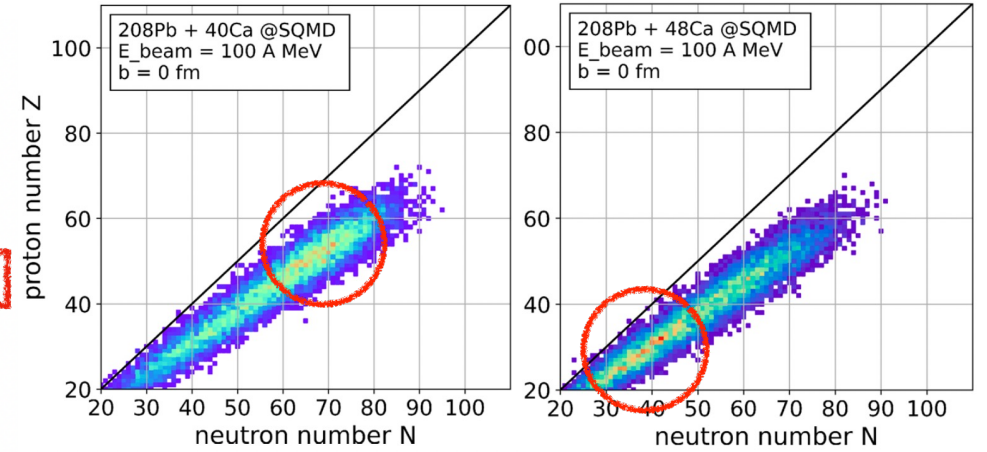
The BFs in DJBUU and SQMD; the BFs from the ten runs of DJBUU and the most abundantly produced three BFs from SQMD runs

Comparative study : Fragment

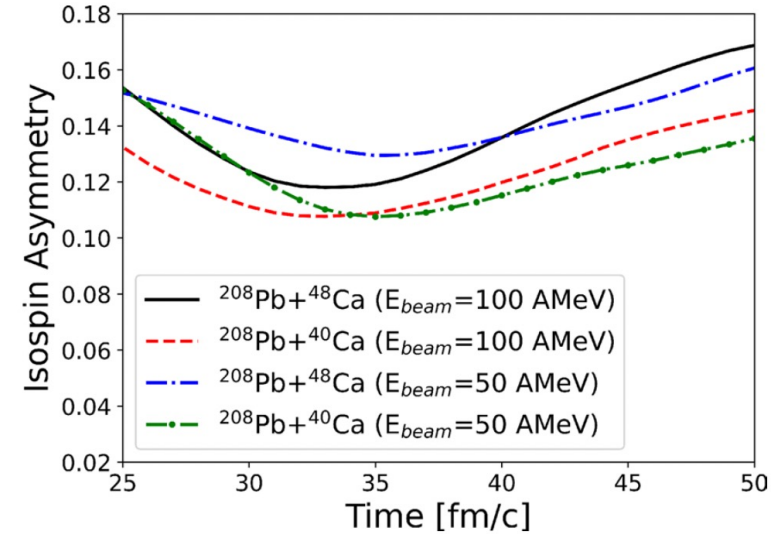


Comparative study : Fragment

| Target | E_{beam} (AMeV) | b (fm) | DJBUU | SQMD |
|------------------|--------------------------|----------|--|--|
| ^{40}Ca | 50 | 0 | $^{163}_{73}\text{Ta}$, $^{162}_{73}\text{Ta}$, $^{164}_{73}\text{Ta}$, $^{163}_{74}\text{W}$ | $^{163}_{69}\text{Tm}$, $^{173}_{74}\text{W}$, $^{169}_{72}\text{Hf}$ |
| | | 3 | $^{163}_{73}\text{Ta}$, $^{165}_{74}\text{W}$, $^{164}_{73}\text{Ta}$ | $^{169}_{72}\text{Hf}$, $^{173}_{74}\text{W}$, $^{172}_{74}\text{W}$ |
| | | 6 | $^{167}_{74}\text{W}$, $^{169}_{75}\text{Re}$, $^{165}_{73}\text{Ta}$, $^{168}_{75}\text{Re}$ | $^{168}_{72}\text{Hf}$, $^{164}_{70}\text{Yb}$, $^{169}_{72}\text{Hf}$ |
| | 100 | 0 | $^{123}_{56}\text{Ba}$, $^{121}_{55}\text{Cs}$, $^{124}_{57}\text{La}$, $^{122}_{56}\text{Ba}$, $^{124}_{56}\text{Ba}$ | $^{78}_{33}\text{As}$, $^{114}_{50}\text{Sn}$, $^{124}_{54}\text{Xe}$ |
| | | 3 | $^{130}_{59}\text{Pr}$, $^{130}_{58}\text{Ce}$, $^{128}_{57}\text{La}$, $^{128}_{58}\text{Ce}$, $^{129}_{58}\text{Ce}$, $^{127}_{58}\text{Ce}$, $^{127}_{57}\text{La}$ | $^{125}_{53}\text{I}$, $^{128}_{56}\text{Ba}$, $^{132}_{57}\text{La}$ |
| | | 6 | $^{145}_{64}\text{Gd}$, $^{144}_{64}\text{Gd}$, $^{146}_{65}\text{Tb}$, $^{147}_{65}\text{Tb}$ | $^{151}_{64}\text{Gd}$, $^{149}_{63}\text{Eu}$, $^{154}_{66}\text{Dy}$ |
| ^{48}Ca | 50 | 0 | $^{161}_{72}\text{Hf}$, $^{162}_{72}\text{Hf}$, $^{160}_{71}\text{Lu}$, $^{159}_{71}\text{Lu}$ | $^{167}_{70}\text{Yb}$, $^{167}_{71}\text{Lu}$, $^{170}_{71}\text{Lu}$ |
| | | 3 | $^{162}_{72}\text{Hf}$, $^{164}_{73}\text{Ta}$ | $^{165}_{70}\text{Yb}$, $^{167}_{70}\text{Yb}$, $^{167}_{71}\text{Lu}$ |
| | | 6 | $^{164}_{72}\text{Hf}$, $^{163}_{72}\text{Hf}$, $^{166}_{73}\text{Ta}$, $^{165}_{72}\text{Hf}$ | $^{165}_{69}\text{Tm}$, $^{159}_{68}\text{Er}$, $^{164}_{69}\text{Tm}$ |
| | 100 | 0 | $^{113}_{51}\text{Sb}$, $^{115}_{52}\text{Te}$, $^{114}_{51}\text{Sb}$, $^{116}_{52}\text{Te}$, $^{112}_{51}\text{Sb}$ | $^{58}_{25}\text{Mn}$, $^{74}_{32}\text{Ge}$, $^{107}_{48}\text{Pd}$ |
| | | 3 | $^{121}_{54}\text{Xe}$, $^{122}_{55}\text{Cs}$, $^{120}_{54}\text{Xe}$, $^{123}_{55}\text{Cs}$, $^{121}_{55}\text{Cs}$ | $^{120}_{52}\text{Te}$, $^{106}_{45}\text{Rh}$, $^{113}_{48}\text{Cd}$ |
| | | 6 | $^{140}_{62}\text{Sm}$, $^{139}_{62}\text{Sm}$, $^{138}_{61}\text{Pm}$, $^{137}_{61}\text{Pm}$, $^{137}_{60}\text{Nd}$ | $^{147}_{62}\text{Sm}$, $^{153}_{64}\text{Gd}$, $^{148}_{62}\text{Sm}$ |



Proton and neutron distributions of the BF in SQMD



The BFs in DJBUU and SQMD; the BFs from the ten runs of DJBUU and the most abundantly produced three BFs from SQMD runs

the symmetry energy pushes out the neutrons and so disturbs the formation of large fragments.

Surface term

Restoring Surface term

Ignore time derivative term and ~~spatial derivative term~~

We need to solve
Numerically
And Efficiently



$$-\nabla^2 \sigma + m^2 \sigma + g_2 \sigma^2 + g_3 \sigma^3 = -g_\sigma \rho_s$$

We found
Analytic solution
for ω, ρ



$$\begin{aligned} -\nabla^2 \omega^0 + m^2 \omega^0 &= g_\omega \rho_B \\ -\nabla^2 \rho^0 + m^2 \rho^0 &= g_\omega \rho_I \\ -\nabla^2 A^0 &= e \rho_q \end{aligned}$$

$$\rho_s = \frac{g}{(2\pi)^3} \sum \int m^* dp^3 f(\vec{x}, \vec{p})$$

- Found Analytic solution for omega, rho fields.
- Couldn't find Analytic solution for sigma fields. → Numerical method, (Jacobi method)

Restoring Surface term

$$-\nabla^2 A^0 = e\rho_q$$

- We have solved Poisson equations
- for Coulomb interaction, with Green's function

3.2 Coulomb Integral

$$\phi_i(\mathbf{x}) = \int d^3x' \frac{1}{4\pi|\mathbf{x} - \mathbf{x}'|} W(\mathbf{x}' - \mathbf{x}_i)$$

...

$$\phi_i(\mathbf{x}) = \theta(a > s) \frac{315}{64\pi a^3} \left(\frac{s^8}{72a^6} - \frac{s^6}{14a^4} + \frac{3s^4}{20a^2} + \frac{a^2}{8} - \frac{s^2}{6} \right) + \theta(s > a) \frac{1}{4\pi s}$$

$$-\nabla^2 \omega^0 + m^2 \omega^0 = g_\omega \rho_B$$

3.3 Yukawa Integral

Courtesy of Prof. Jeon

$$\phi_i(\mathbf{x}) = \int d^3x' \frac{e^{-m|\mathbf{x}-\mathbf{x}'|}}{4\pi|\mathbf{x} - \mathbf{x}'|} W(\mathbf{x}' - \mathbf{x}_i) \quad (43)$$

...

Defining $ma = \tilde{a}$ and $ms = \tilde{s}$,

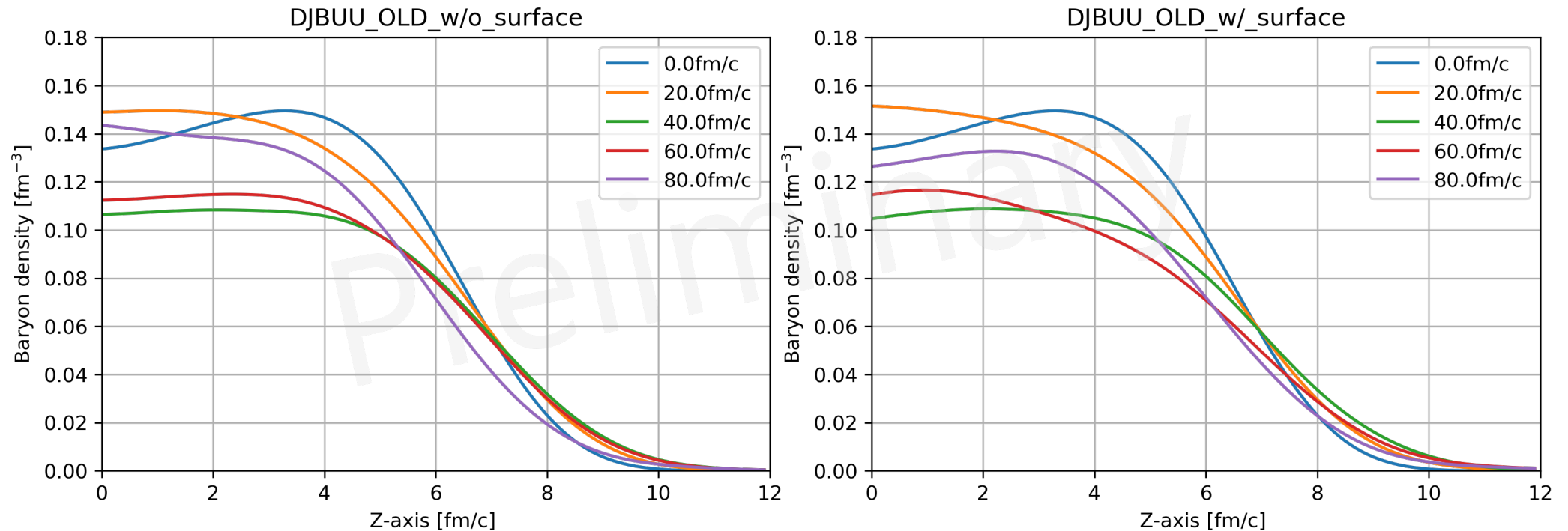
$\phi_i(\mathbf{x})$

$$= \theta(s < a) \frac{315m}{64\pi\tilde{a}^9\tilde{s}}$$

$$\left(48(\tilde{a}(\tilde{a}(\tilde{a}(\tilde{a} + 10) + 45) + 105) + 105) \sinh(\tilde{s})e^{-\tilde{a}} - \tilde{s}(-\tilde{a}^6 + 3\tilde{a}^4(\tilde{s}^2 + 6) - 3\tilde{a}^2(\tilde{s}^4 + 20\tilde{s}^2 + 120) + 840(\tilde{s}^2 + 6) + \tilde{s}^4(\tilde{s}^2 + 42)) \right) + \theta(s > a) \frac{945me^{-\tilde{s}}}{4\pi\tilde{a}^9\tilde{s}} \left((\tilde{a}^4 + 45\tilde{a}^2 + 105) \sinh(\tilde{a}) - 5\tilde{a}(2\tilde{a}^2 + 21) \cosh(\tilde{a}) \right) \quad (54)$$

Stability (w/o and w/ Surface term)

$^{197}\text{Au} + ^{197}\text{Au}$, $E_{beam} = 50 \text{ A MeV}$, $b = 40.0 \text{ fm}$

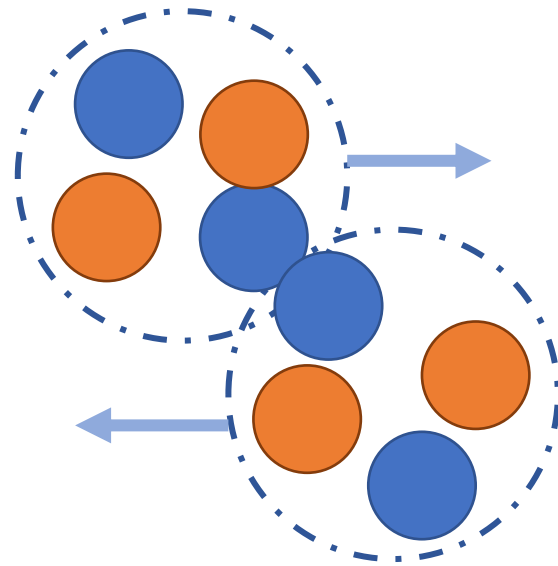


^{197}Au Stability in the DJBUU simulation w/o surface term (left), w/ surface term (right)

QMC model

Adopting Quark-Meson Coupling model

Quantum Hadron Dynamics (QHD)
Such as Walecka model



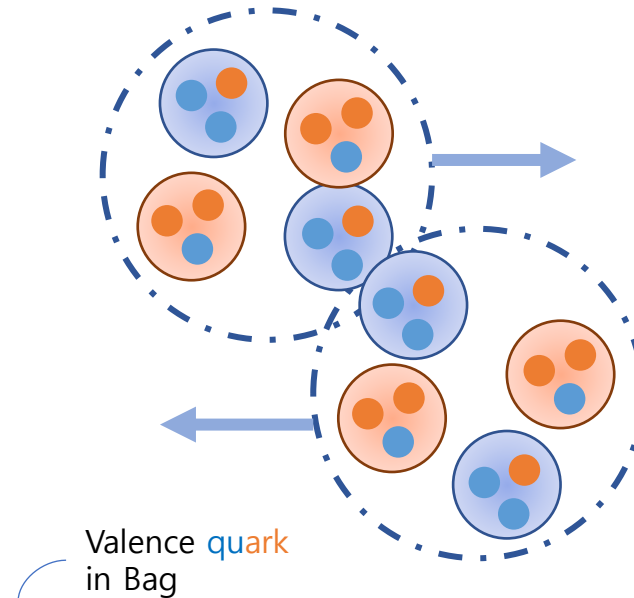
Nucleon-meson coupling

$$g_{\sigma}^N, g_{\omega}^N, g_{\rho}^N$$

Effective
mass

$$m^* = m - g\sigma$$

Quark-Meson Coupling (QMC)



Quark-meson coupling

$$g_{\sigma}^{u,d}, g_{\sigma}^s \rightarrow g_{\sigma}^{N,Y} \dots$$

$$m^* = m - g\sigma + \frac{d}{2}(g\sigma)^2$$

Scalar polarization

Adopting Quark-Meson Coupling model

Lagrangian for quark

$$\mathcal{L} = [\bar{\psi}(i\partial - m_q - g_\sigma^q \sigma - g_\omega^q \omega - g_\rho^q \boldsymbol{\tau} \cdot \boldsymbol{\rho})\psi - B]\theta_V - \frac{1}{2}\bar{\psi}\psi\theta_S$$

Lagrangian for nucleon

$$\mathcal{L} = \bar{\psi}[i\partial - (m_N - g_\sigma(\sigma)) - g_\omega \omega - g_\rho \boldsymbol{\tau} \cdot \boldsymbol{\rho}]\psi$$

$$m_N - g\sigma \rightarrow m_N - (g\sigma - \frac{a_N}{2}(g\sigma)^2)$$

Simple parameterization

$$g_\sigma(\sigma) = g_{\sigma=0} - \frac{a_N}{2} g_{\sigma=0}^2 \sigma$$

$$C_N(\sigma) = 1 - a_N g_{\sigma=0} \sigma$$

Meson eqs. in QHD

$$-\nabla^2 \sigma + m^2 \sigma + g_2 \sigma^2 + g_3 \sigma^3 = -g_\sigma \rho_S$$

$$-\nabla^2 \omega^0 + m^2 \omega^0 = g_\omega \rho_B$$

$$-\nabla^2 \rho^0 + m^2 \rho^0 = g_\omega \rho_I$$



Meson eqs. in QMC

$$\omega = \frac{g_\omega}{m_\omega^2} \rho_B \equiv \frac{g_\omega}{m_\omega^2} \frac{4}{(2\pi)^3} \int d^3k \theta(k_F - |\vec{k}|),$$

$$\sigma = \frac{g_\sigma^N}{m_\sigma^2} C_N(\sigma) \rho_S \equiv \frac{g_\sigma^N}{m_\sigma^2} C_N(\sigma) \frac{4}{(2\pi)^3} \int d^3k \theta(k_F - |\vec{k}|) \frac{m_N^*(\sigma)}{\sqrt{m_N^{*2}(\sigma) + \vec{k}^2}},$$

PTEP, 2022 043D02

Adopting Quark-Meson Coupling model

$$m_B^* \simeq m_B - \frac{n_q}{3} g_\sigma^N \left[1 - \frac{a_B}{2} (g_\sigma^N \sigma) \right] \sigma = m_B - \frac{n_q}{3} \left[(g_\sigma^N \sigma) - \frac{a_B}{2} (g_\sigma^N \sigma)^2 \right]$$

scalar polarizability

($B = N, \Lambda, \Sigma, \Xi, \Delta, \Sigma^*, \Xi^*, \Lambda_c, \Sigma_c, \Xi_c, \Lambda_b, \Sigma_b, \Xi_b$),

$$\sigma = \frac{g_\sigma^N}{m_\sigma^2} C_N(\sigma) \rho_s \equiv \frac{g_\sigma^N}{m_\sigma^2} C_N(\sigma) \frac{4}{(2\pi)^3} \int d^3k \theta(k_F - |\vec{k}|) \frac{m_N^*(\sigma)}{\sqrt{m_N^{*2}(\sigma) + \vec{k}^2}},$$

PTEP, 2022 043D02

Density Parameterization, Prof. Tsushima

($\rho_0 = 0.15 \text{ fm}^{-3}$) as

$$(g_\sigma^N \sigma)(x) = \begin{cases} 1.60828 - 23.9107\sqrt{x} + 350.631x & (x > 0), \\ -144.309x\sqrt{x} + 19.4750x^2 & (x > 0), \\ 0 & (x = 0), \end{cases} \quad (20)$$

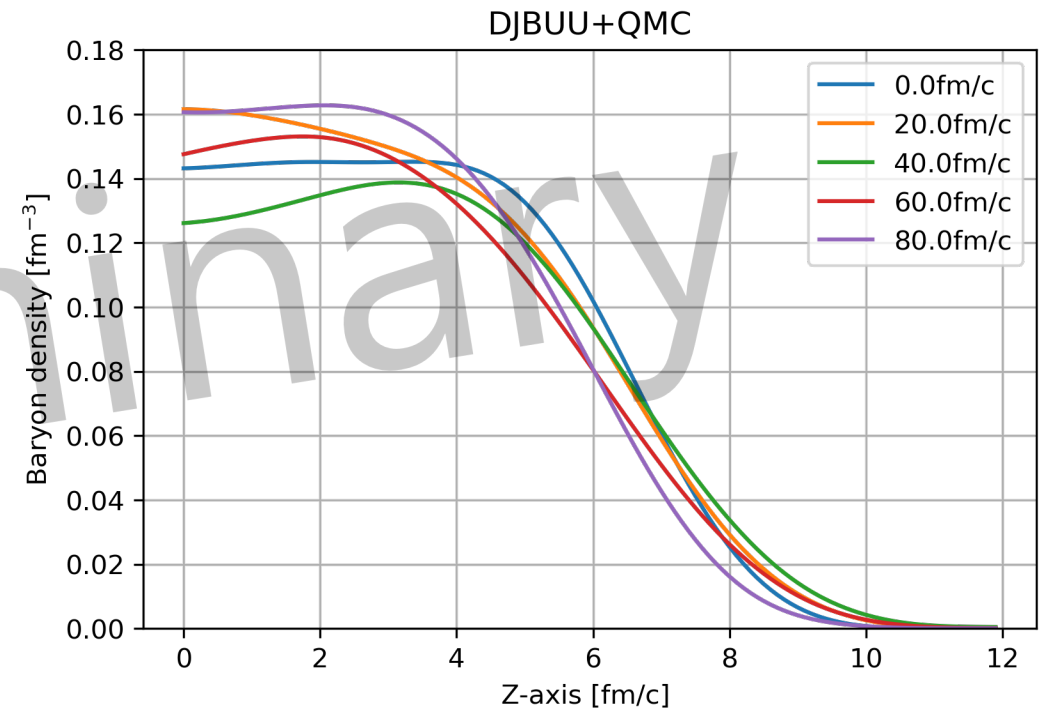
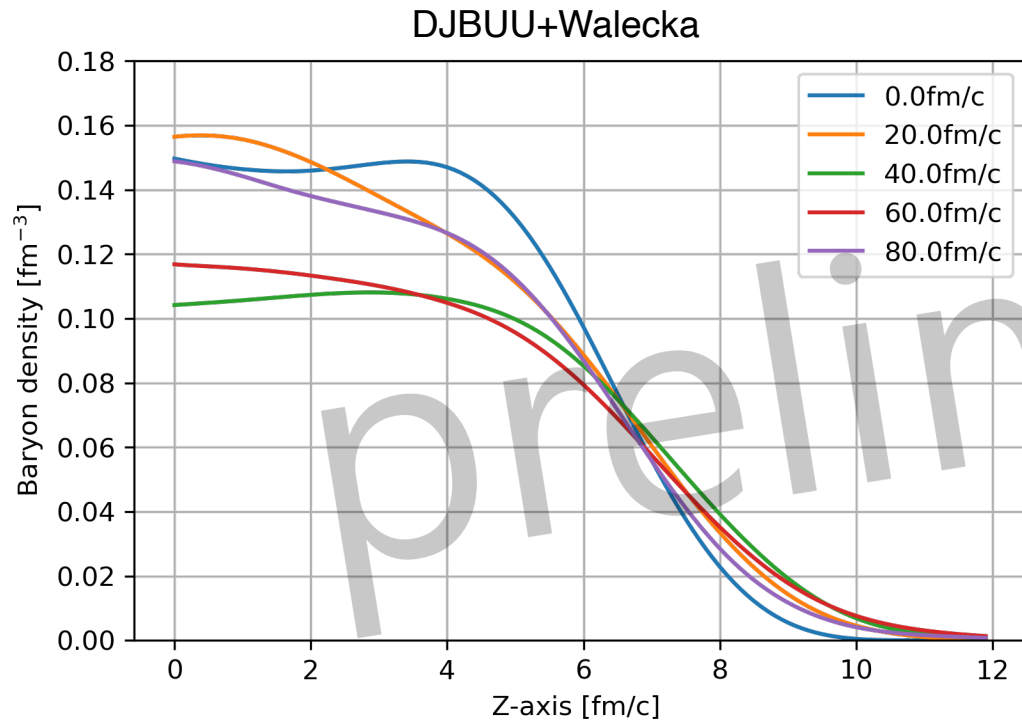
$$V_\omega^B(x) = b_B x,$$

For completeness, we also give the Lorentz-vector-isovector mean field potential (in MeV) as a function of $y \equiv \rho_3/\rho_0 = (\rho_p - \rho_n)/\rho_0$ with the isospin-third component of the hadron h , I_3^h ,

$$I_3^h V_\rho^h(y) = I_3^h \times 84.61y, \quad (24)$$

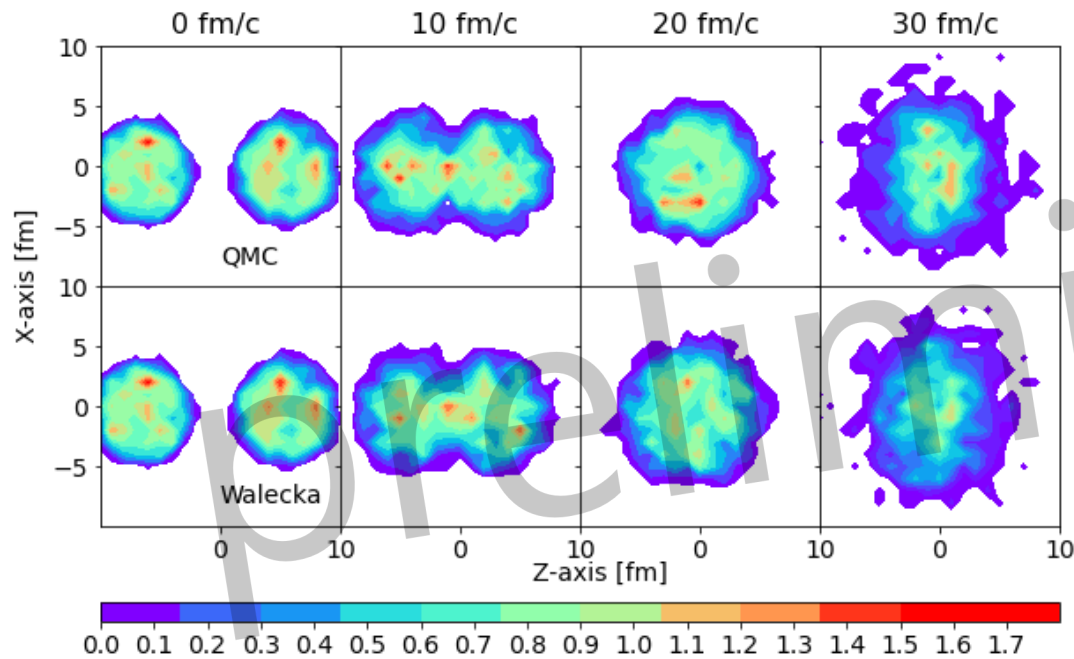
Adopting Quark-Meson Coupling model

Check stability $^{197}\text{Au} + ^{197}\text{Au}$, $E_{\text{beam}} = 50 \text{ A MeV}$

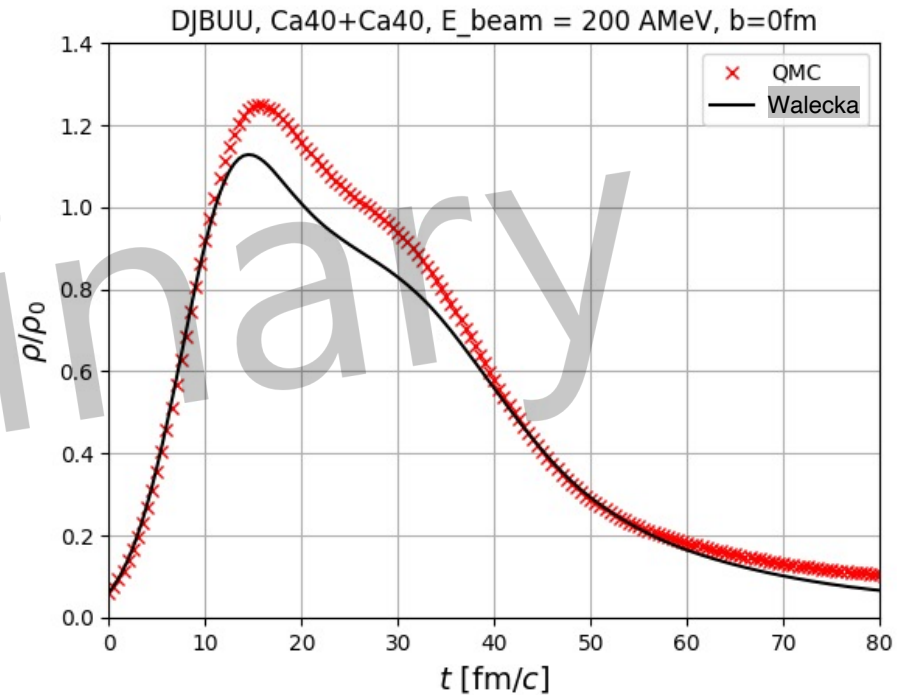


Adopting Quark-Meson Coupling model

$$^{40}\text{Ca} + ^{40}\text{Ca}, E_{\text{beam}} = 200 \text{ A MeV}, b = 0 \text{ fm}$$



Contour QMC (top), QHD-Walecka (bottom)



Baryon density at center of mass

- Central density with QMC is higher than one with QHD

Adopting Quark-Meson Coupling model

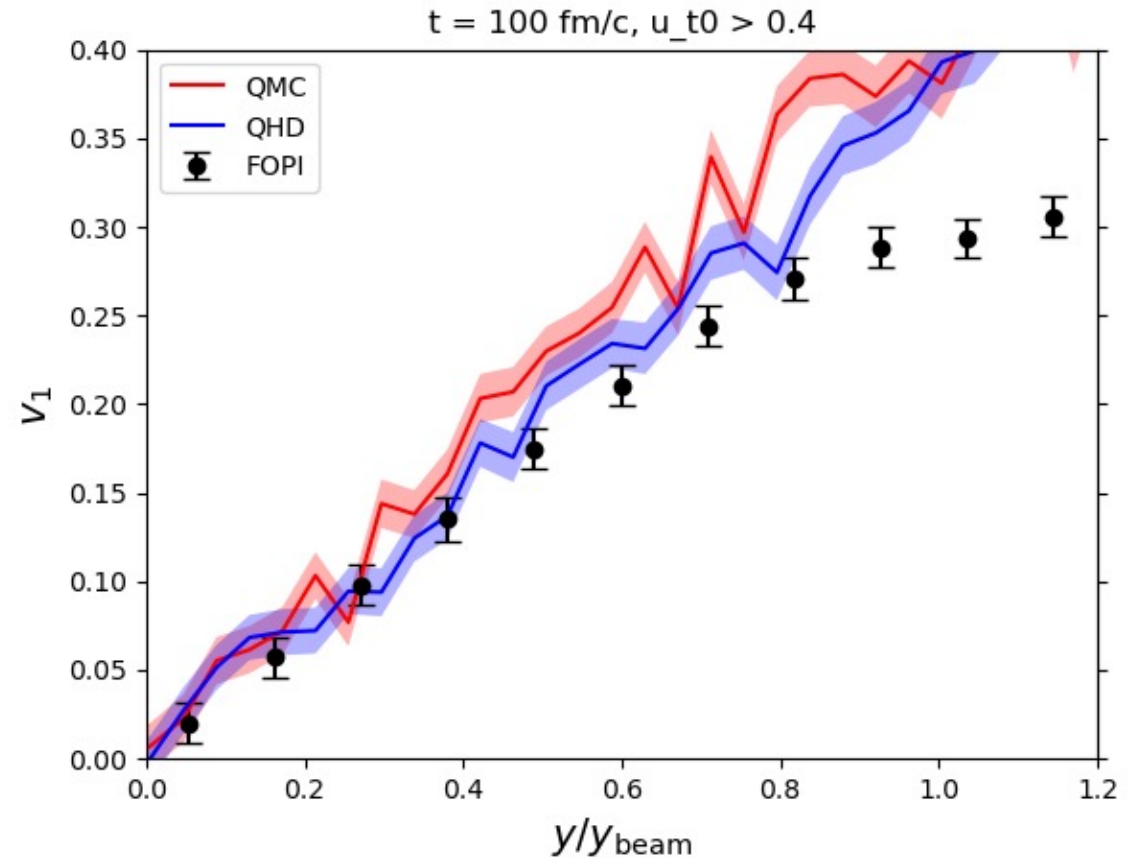
$^{197}\text{Au} + ^{197}\text{Au}$, $E_{\text{beam}} = 400 \text{ A MeV}$, $0.25 < b_0 < 0.45$

$$b_0 = 1.15 \times (A_{\text{projectile}}^{1/3} + A_{\text{target}}^{1/3})$$

$$b_0 = 0.25 \rightarrow b = 3.346 \text{ fm}$$

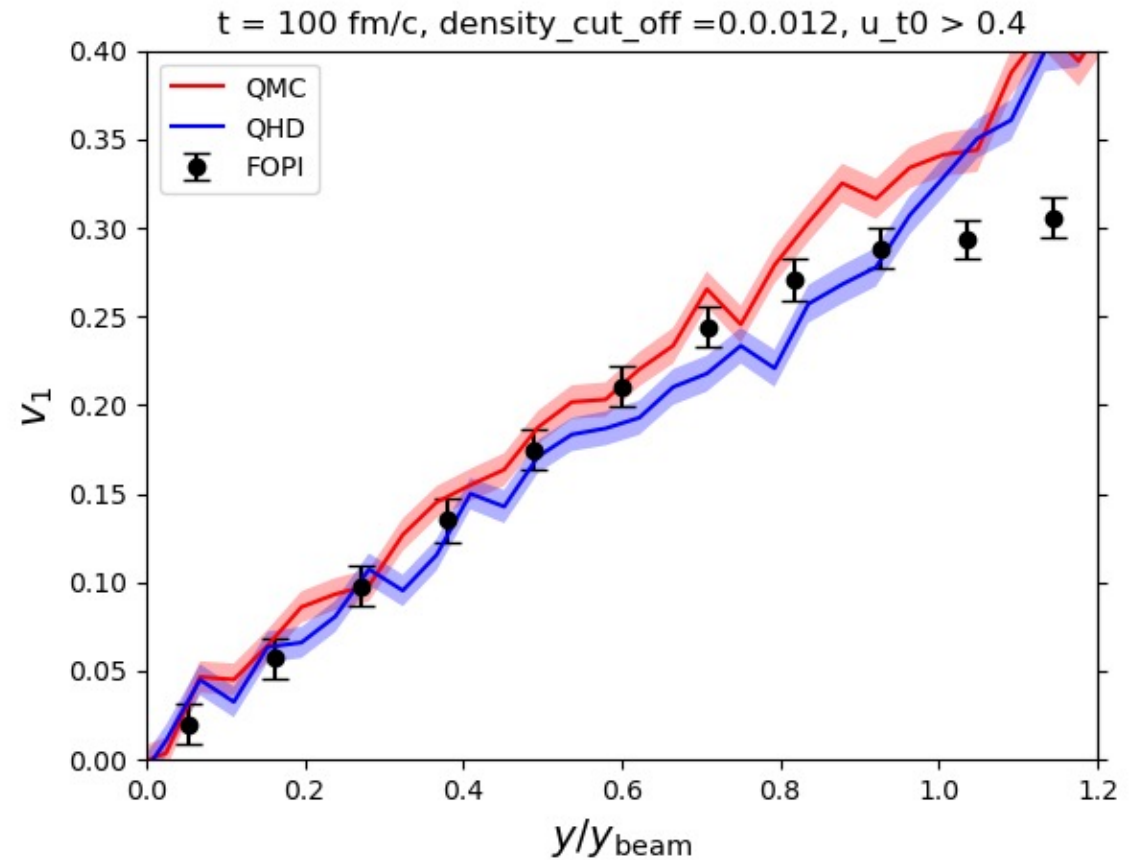
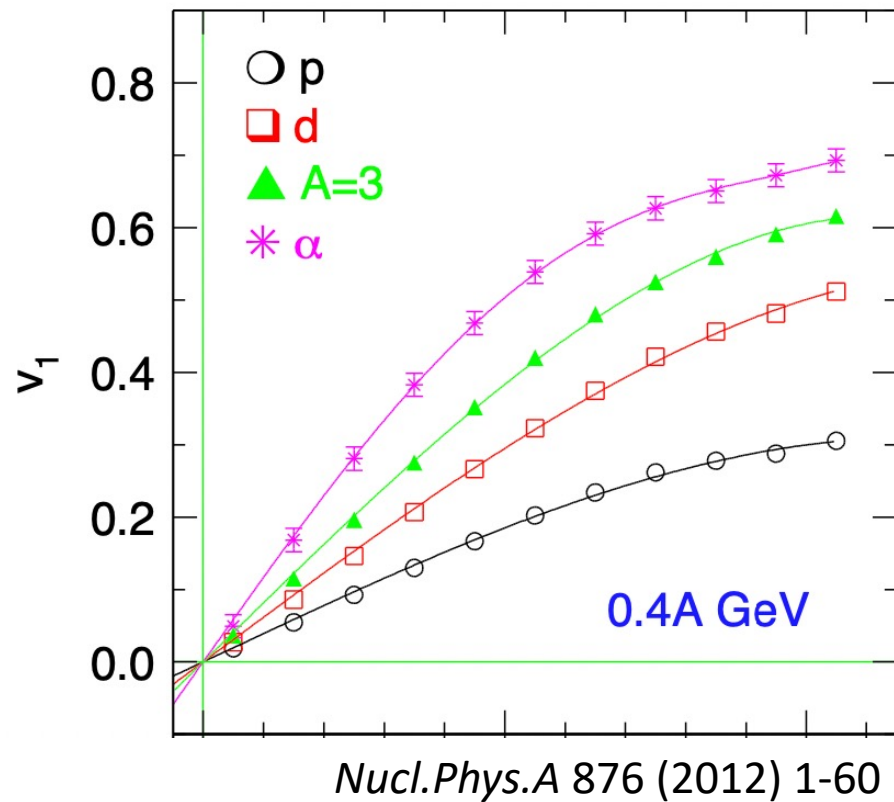
$$b_0 = 0.35 \rightarrow b = 4.684 \text{ fm}$$

$$b_0 = 0.45 \rightarrow b = 6.022 \text{ fm}$$



Adopting Quark-Meson Coupling model

$^{197}\text{Au} + ^{197}\text{Au}$, $E_{\text{beam}} = 400 \text{ A MeV}$, $0.25 < b_0 < 0.45$



Thank you for your attention

Any questions?

Acknowledgement

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BACK UP SLICES